

Utilizing Symmetry Coupling Schemes in *Ab Initio* Nuclear Structure Calculations

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Abstract

We report on *ab initio* no-core shell model calculations in a symmetry-adapted SU(3)-based coupling scheme that demonstrate that collective modes in *p*-shell nuclei emerge from first principles. The low-lying states of ⁶Li, ⁶He, ⁸Be, ⁸B, ¹²C, and ¹⁶O, are shown to exhibit orderly patterns that favor spatial configurations with strong quadrupole deformation and complementary low intrinsic spin values, a picture that is consistent with the nuclear symplectic model. The results also suggest a pragmatic path forward to accommodate deformation-driven collective features in *ab initio* analyses when they dominate the nuclear landscape.

Keywords: *No-core shell model; SU(3) coupling scheme; p-shell nuclei*

1 Introduction

In the last few years, *ab initio* approaches to nuclear structure and reactions have considerably advanced our understanding and capability of achieving first-principle descriptions of *p*-shell nuclei [1–3]. These advances are driven by the major progress in the development of realistic nuclear potential models, such as *J*-matrix inverse scattering potentials [4] and two- and three-nucleon potentials derived from meson exchange theory [5] or by using chiral effective field theory [6], and, at the same time, by the utilization of massively parallel computing resources [7–9].

The predictive power that *ab initio* models hold [10, 11] makes them suitable for targeting short-lived nuclei that are inaccessible by experiment but essential to further modeling, for example, of the dynamics of X-ray bursts and the path of nucleosynthesis (see, e. g., Refs. [12, 13]). The main limitation of *ab initio* approaches is inherently coupled with the combinatorial growth in the size of the many-particle model space with increasing nucleon numbers and expansion in the number of single-particle levels in the model space as illustrated in Fig. 1. This points to the need of further major advances in many-body methods to access a wider range of nuclei and experimental observables, while retaining the *ab initio* predictive power.

These considerations motivate us to develop and investigate a novel model, the *ab initio* symmetry-adapted no-core shell model (SA-NCSM) [14], which by taking advantage of symmetries inherent to the nuclear dynamics [15, 16] allows one to truncate a model space according to correlations indispensable for modeling important modes of nuclear collective dynamics, thereby overcoming the scale explosion bottleneck of *ab initio* nuclear structure computations.

Proceedings of International Conference ‘Nuclear Theory in the Supercomputing Era — 2013’ (NTSE-2013), Ames, IA, USA, May 13–17, 2013. Eds. A. M. Shirokov and A. I. Mazur. Pacific National University, Khabarovsk, Russia, 2014, p. 62.

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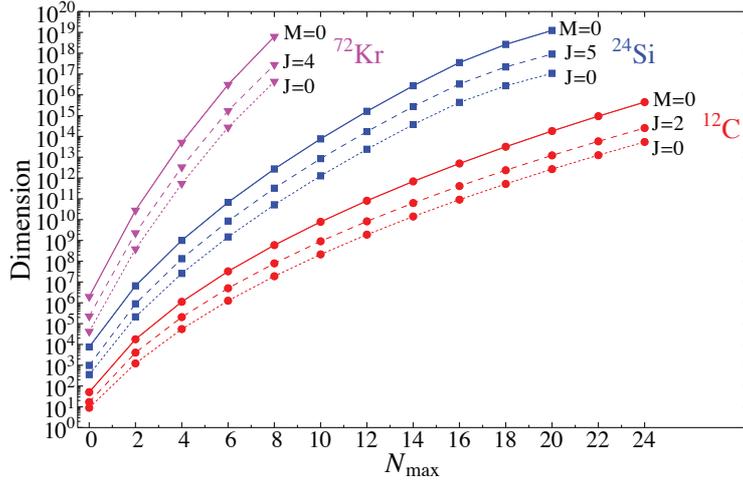


Figure 1: The dimensions of positive parity model spaces as functions of N_{\max} for selected nuclei. Solid curves show the number of basis states with the projection of the total angular momentum $M = 0$. Dashed and dotted curves depict the number of basis states carrying selected values of the total angular momentum J .

2 *Ab initio* calculations in a SU(3)-based coupling scheme

The SA-NCSM joins a no-core shell model (NCSM) theory [2] with a multi-shell, SU(3)-based coupling scheme [15, 17]. Specifically, the many-nucleon basis states of the SA-NCSM are decomposed into spatial and intrinsic spin parts, where the spatial part is further classified according to the $SU(3) \supset SO(3)$ group chain. The significance of the SU(3) group for a microscopic description of the nuclear collective dynamics can be seen from the fact that it is the symmetry group of the successful Elliott model [15], and a subgroup of the physically relevant $Sp(3, \mathbb{R})$ symplectic model [16], which provides a comprehensive theoretical foundation for understanding the dominant symmetries of nuclear collective motion. The SA-NCSM basis states are labeled as

$$|\bar{\gamma}; N(\lambda\mu)\kappa L; (S_p S_n)S; JM\rangle, \quad (1)$$

where N signifies the number of harmonic oscillator quanta with respect to the minimal number for a given nucleus. Quantum numbers S_p , S_n , and S denote proton, neutron, and total intrinsic spins, respectively, and $(\lambda\mu)$ represent a set of quantum numbers associated with SU(3) irreducible representations, irreps. The label κ distinguishes multiple occurrences of the same orbital momentum L in the parent irrep $(\lambda\mu)$. The L is coupled with S to the total angular momentum J and its projection M . The basis states bring forward important information about nuclear shapes and deformation according to an established mapping [18]. For example, (00) , $(\lambda 0)$ and (0μ) describe spherical, prolate and oblate shapes, respectively. The symbol $\bar{\gamma}$ schematically denotes the additional quantum numbers needed to specify a distribution of nucleon clusters over the major HO shells and their inter-shell coupling. Specifically, in each major HO shell η with degeneracy Ω_η , clusters of protons and neutrons are arranged into antisymmetric $U(\Omega_\eta) \times SU(2)_{S_\eta}$ irreps [19], with $U(\Omega_\eta)$ further reduced with respect to SU(3). The quantum numbers, $[f_1, \dots, f_{\Omega_\eta}] \alpha_\eta (\lambda_\eta \mu_\eta) S_\eta$, along with $SU(3) \times SU(2)_S$ labels of inter-shell coupling unambiguously determine SA-NCSM basis states (1). Note that a spatial symmetry associated with a Young

shape $[f_1, \dots, f_{\Omega_\eta}]$ is uniquely determined by the imposed antisymmetrization and the associated intrinsic spin S_η . A multiplicity index α_η is required to distinguish multiple occurrences of SU(3) irrep $(\lambda_\eta \mu_\eta)$ in a given $U(\Omega_\eta)$ irrep. It is important to note that any model space spanned by a complete set of equivalent $SU(3) \times SU(2)_S$ irreps, that is, a space spanned by all configurations carrying a fixed set of $S_p S_n S$ and $(\lambda \mu)$ quantum numbers, permits exact factorization of the center-of-mass motion.

The SA-NCSM implements fast methods for calculating matrix elements of arbitrary (currently up to two-body, but expandable to higher-rank) operators in the symmetry-adapted basis. This facilitates both the evaluation of the Hamiltonian matrix elements and the use of the resulting eigenvectors to evaluate other experimental observables. The underlying principle behind the SA-NCSM computational kernel is an SU(3)-type Wigner–Eckhart theorem, which factorizes interaction matrix elements into the product of SU(3) reduced matrix elements (*rme*) and the associated SU(3) coupling coefficient. The SA-NCSM configurations are constructed by the inter-shell coupling of a sequence of single-shell nucleon clusters arranged into $U(\Omega) \times SU(2)_S$, with $U(\Omega) \supset SU(3)$, irreps. Therefore, all the multi-shell *rme* are constructed from a set of single-shell *rme* computed in a configuration space of these irreps. This reduces the number of key pieces of information required to the single-shell *rme*, and these track with the number of $U(\Omega) \times SU(2)_S$ irreps, with $U(\Omega) \supset SU(3)$, that represent building blocks of the SA-NCSM approach. It is therefore significant that their number grows slowly with the increasing nucleon number and N_{\max} cutoff as this allows these key pieces of information to be stored in CPU memory.

3 Structure of nuclear wave functions

The expansion of calculated eigenstates in the physically relevant SU(3) basis unveils salient features that emerge from the complex dynamics of these strongly interacting many-particle systems. To explore the nature of the most important correlations, we analyze the probability distribution across Pauli-allowed $(S_p S_n S)$ and $(\lambda \mu)$ configurations of the four lowest-lying isospin-zero ($T = 0$) states of ${}^6\text{Li}$ (1_{gs}^+ , 3_1^+ , 2_1^+ , and 1_2^+), the ground-state rotational bands of ${}^8\text{Be}$, ${}^6\text{He}$ and ${}^{12}\text{C}$, the lowest 1^+ , 3^+ , and 0^+ excited states of ${}^8\text{B}$, and the ground state of ${}^{16}\text{O}$. Results for the ground state of ${}^6\text{Li}$ and ${}^8\text{Be}$, obtained with the JISP16 and chiral $N^3\text{LO}$ interactions, respectively, are shown in Figs. 2 and 3. These figures illustrate a feature common to all the low-energy solutions considered; namely, a highly structured and regular mix of intrinsic spins and SU(3) spatial quantum numbers that has heretofore gone unrecognized in other *ab initio* studies, and which, furthermore, does not seem to depend on the particular choice of realistic NN potential.

For a closer look at these results, first consider the spin content. We found that the calculated eigenstates project at a 99% level onto a comparatively small subset of intrinsic spin combinations. For instance, the lowest-lying eigenstates in ${}^6\text{Li}$ are almost entirely realized in terms of configurations characterized by the following intrinsic spin $(S_p S_n S)$ triplets: $(\frac{3}{2} \frac{3}{2} 3)$, $(\frac{1}{2} \frac{3}{2} 2)$, $(\frac{3}{2} \frac{1}{2} 2)$, and $(\frac{1}{2} \frac{1}{2} 1)$, with the last one carrying over 90% of each eigenstate. Likewise, the same spin components as in the case of ${}^6\text{Li}$ are found to dominate the ground state and the lowest 1^+ , 3^+ , and 0^+ excited states of ${}^8\text{B}$ (Table 1). The ground state bands of ${}^8\text{Be}$, ${}^6\text{He}$, ${}^{12}\text{C}$, and ${}^{16}\text{O}$ are found to be dominated by many-particle configurations carrying total intrinsic spin of the protons and neutrons equal to zero and one, with the largest contributions due to $(S_p S_n S) = (000)$ and (112) configurations.

Second, consider the spatial degrees of freedom. Our results show that the mixing of $(\lambda \mu)$ quantum numbers, induced by the SU(3) symmetry breaking terms of realistic interactions, exhibits a remarkably simple pattern. One of its key features is the preponderance of a single $0\hbar\Omega$ SU(3) irrep. This so-called leading irrep, according to the established geometrical interpretation of SU(3) labels $(\lambda \mu)$ [18], is characterized

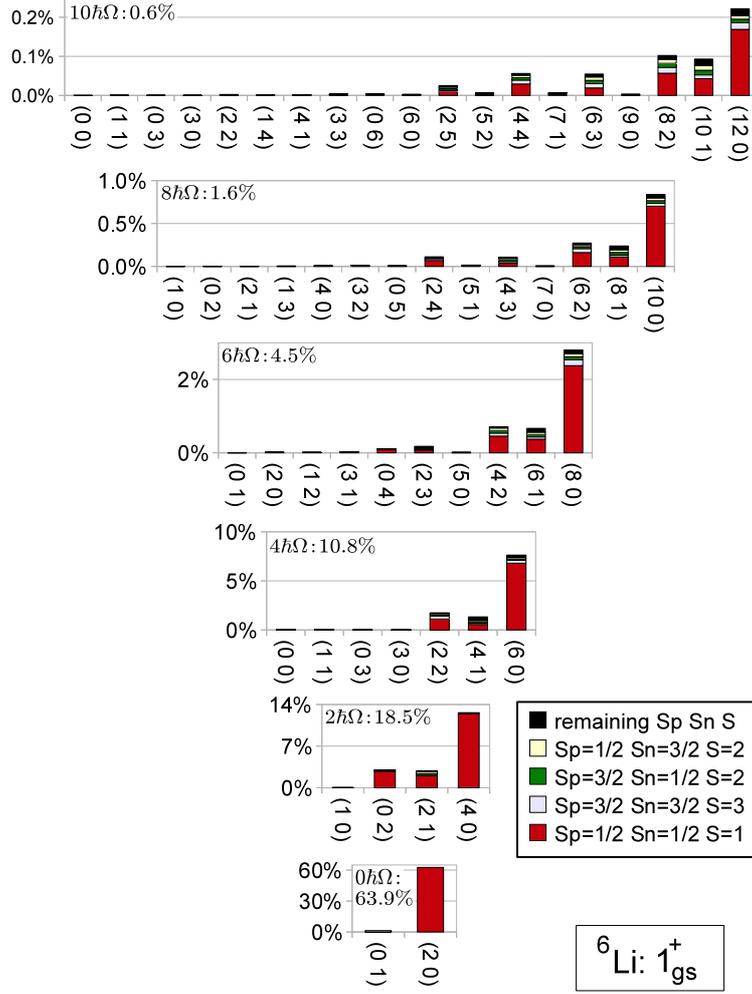


Figure 2: Probability distributions for proton, neutron, and total intrinsic spin components ($S_p S_n S$) across the Pauli-allowed $(\lambda \mu)$ values (horizontal axis) for the calculated 1^+ ground state of ${}^6\text{Li}$ obtained for $N_{\text{max}} = 10$ and $\hbar\Omega = 20$ MeV with the JISP16 interaction. The total probability for each $N\hbar\Omega$ subspace is given in the upper left-hand corner of each histogram. Adapted from Ref. [14].

by the largest value of the intrinsic quadrupole deformation. For instance, the low-lying states of ${}^6\text{Li}$ project at a 40%–70% level onto the prolate $0\hbar\Omega$ SU(3) irrep (20), as illustrated in Figs. 2 and 3 for the ground state. For the considered states of

Table 1: Probability amplitude of the dominant ($S_p S_n S$) spin configuration and the dominant nuclear shapes according to Eq. (2) for the ground state of p -shell nuclei.

Nucleus	$(S_p S_n S)$	Prob. [%]	$(\lambda_0 \mu_0)$	Prob. [%]
${}^6\text{Li}$	$(\frac{1}{2} \frac{1}{2} 1)$	93.26	(20)	98.13
${}^8\text{B}$	$(\frac{1}{2} \frac{1}{2} 1)$	85.17	(21)	87.94
${}^8\text{Be}$	(000)	85.25	(40)	90.03
${}^{12}\text{C}$	(000)	55.19	(04)	48.44
${}^{16}\text{O}$	(000)	83.60	(00)	89.51

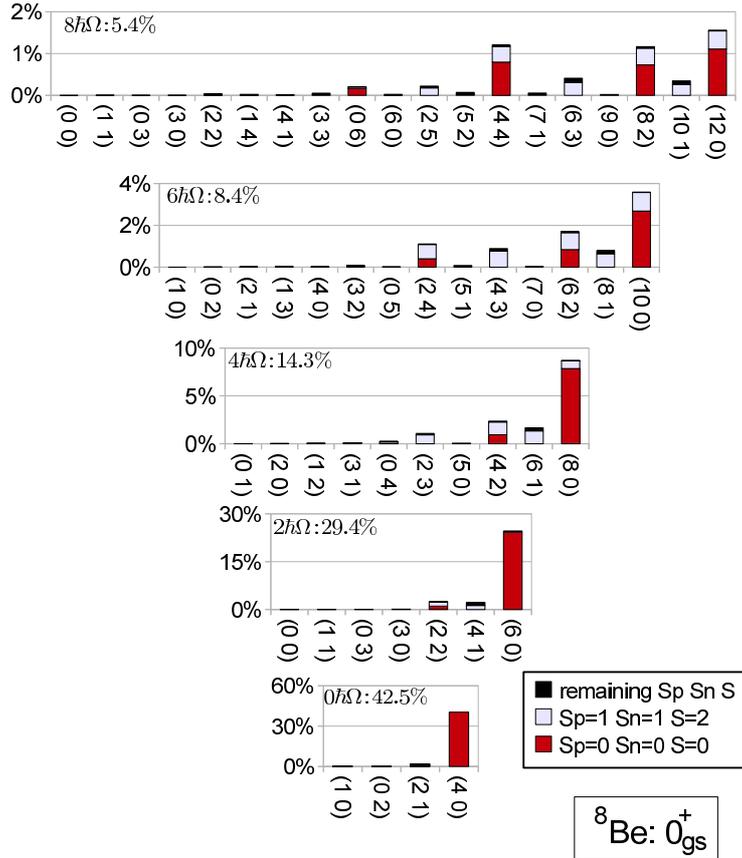


Figure 3: Probability distributions for proton, neutron, and total intrinsic spin components ($S_p S_n S$) across the Pauli-allowed $(\lambda\mu)$ values (horizontal axis) for the calculated 0^+ ground state of ${}^8\text{Be}$ obtained for $N_{\text{max}} = 8$ and $\hbar\Omega = 25$ MeV with the chiral N^3LO interaction. The total probability for each $N\hbar\Omega$ subspace is given in the upper left-hand corner of each histogram. Adapted from Ref. [14].

${}^8\text{B}$, ${}^8\text{Be}$, ${}^{12}\text{C}$, and ${}^{16}\text{O}$, qualitatively similar dominance of the leading $0\hbar\Omega$ $\text{SU}(3)$ irreps is observed — $(2\ 1)$, $(4\ 0)$, $(0\ 4)$, and $(0\ 0)$ irreps, associated with triaxial, prolate, oblate, and spherical shapes, respectively. The clear dominance of the most deformed $0\hbar\Omega$ configuration within low-lying states of light p -shell nuclei indicates that the quadrupole-quadrupole interaction of the Elliott $\text{SU}(3)$ model of nuclear rotations [15] is realized naturally within an *ab initio* framework.

The analysis also reveals that the dominant $\text{SU}(3)$ basis states at each $N\hbar\Omega$ subspace ($N = 0, 2, 4, \dots$) are typically those with $(\lambda\mu)$ quantum numbers given by

$$\lambda + 2\mu = \lambda_0 + 2\mu_0 + N \quad (2)$$

where λ_0 and μ_0 denote labels of the leading $\text{SU}(3)$ irrep in the $0\hbar\Omega$ ($N = 0$) subspace. We conjecture that this regular pattern of $\text{SU}(3)$ quantum numbers reflects the presence of an underlying symplectic $\text{Sp}(3, \mathbb{R})$ symmetry of microscopic nuclear collective motion [16] that governs the low-energy structure of both even-even and odd-odd p -shell nuclei. This can be seen from the fact that $(\lambda\mu)$ configurations that satisfy condition (2) can be determined from the leading $\text{SU}(3)$ irrep $(\lambda_0\ \mu_0)$ through a successive application of a specific subset of the $\text{Sp}(3, \mathbb{R})$ symplectic $2\hbar\Omega$ raising operators. This subset is composed of the three operators, \hat{A}_{zz} , \hat{A}_{zx} , and \hat{A}_{xx} , that distribute two oscillator quanta in z and x directions, but none in y direction, thereby inducing

SU(3) configurations with ever-increasing intrinsic quadrupole deformation. These three operators are the generators of the $\text{Sp}(2, \mathbb{R}) \subset \text{Sp}(3, \mathbb{R})$ subgroup [20], and give rise to deformed shapes that are energetically favored by an attractive quadrupole-quadrupole interaction [21]. Note that this is consistent with our earlier findings of a clear symplectic $\text{Sp}(3, \mathbb{R})$ structure with the same pattern (2) in *ab initio* eigensolutions for ^{12}C and ^{16}O [22].

Furthermore, there is an apparent hierarchy among states that fulfill condition (2). In particular, the $N\hbar\Omega$ configurations with $(\lambda_0 + N \mu_0)$, the so-called stretched states, carry a noticeably higher probability than the others. For instance, the $(2 + N 0)$ stretched states contribute at the 85% level to the ground state of ^6Li , as can be readily seen in Figs. 2 and 3. Moreover, the dominance of the stretched states is rapidly increasing with the increasing many-body basis cutoff N_{max} as illustrated in Fig. 4. The sequence of the stretched states is formed by consecutive applications of the \hat{A}_{zz} operator, the generator of $\text{Sp}(1, \mathbb{R}) \subset \text{Sp}(2, \mathbb{R}) \subset \text{Sp}(3, \mathbb{R})$ subgroup, over the leading SU(3) irrep. This translates into distributing N oscillator quanta along the direction of the z -axis only and hence rendering the largest possible deformation. The important role of the stretched configurations for the description of the rotational bands in $N = Z$ even-even nuclei was recognized heretofore using a simple microscopic Hamiltonian [23]. In the present study, for the first time, this structure is clearly and simply unveiled within the context of a fully microscopic framework starting from first principles.

4 Efficacy of the SU(3) basis

The observed patterns of intrinsic spin and deformation mixing supports a symmetry-guided basis selection philosophy referenced above. Specifically, one can take advantage of dominant symmetries to refine the definition of the NCSM model space, which is based solely on the N_{max} cutoff. A SA-NCSM model space, which we denote as $\langle N_{\text{max}}^\perp \rangle N_{\text{max}}^\top$, can be constructed using a symmetry-guided selection that includes the complete basis up through some $N_{\text{max}}^\perp \leq N_{\text{max}}$ along with configurations carrying a restricted set of $(\lambda \mu)$ and $(S_p S_n S)$ quantum numbers in the N_{max}^\perp to N_{max}^\top space. Ultimately, we aim to achieve $N_{\text{max}}^\top \geq N_{\text{max}}$, where N_{max} is the largest value for which complete-space results can be currently calculated. This concept focuses on retaining the most important configurations that support the strong many-nucleon correlations of a nuclear system using the underlying $\text{Sp}(1, \mathbb{R}) \subset \text{Sp}(2, \mathbb{R}) \subset \text{Sp}(3, \mathbb{R})$ symmetry considerations. Within this context, it is important to note that for model spaces truncated according to $(\lambda \mu)$ irreps and intrinsic spins $(S_p S_n S)$, the spurious center-of-mass motion can be factored out exactly, which represents an important advantage of this scheme.

The efficacy of the symmetry-guided concept is illustrated for SA-NCSM results obtained in a model space, which is expanded beyond the complete $N_{\text{max}}^\perp = 6$ (or 8) space by relatively few dominant intrinsic spin components and quadrupole deformations that satisfy condition (2). We use selected spaces up through $N_{\text{max}}^\top = 12$, which allows a comparison to available results obtained in the complete $N_{\text{max}} = 12$ space and hence, probes the efficacy of the SA-NCSM symmetry-guided model space selection concept. For this analysis, a Coulomb plus bare JISP16 NN interaction for $\hbar\Omega$ values ranging from 17.5 up to 25 MeV is used. SA-NCSM eigenstates are used to determine spectroscopic properties of low-lying $T = 0$ states of ^6Li for a $\langle 6 \rangle 12$ model space and of the ground-state band of ^6He for $\langle 8 \rangle 12$. We utilize a complete space of $N_{\text{max}}^\perp = 6$ for ^6Li and of $N_{\text{max}}^\perp = 8$ for ^6He , as these spaces seem sufficient to accommodate essential mixing of low-energy HO excitations.

The results indicate that the observables obtained in the symmetry-guided truncated spaces under consideration are excellent approximations to the corresponding complete-space counterparts. In particular, the ground-state binding energies repre-

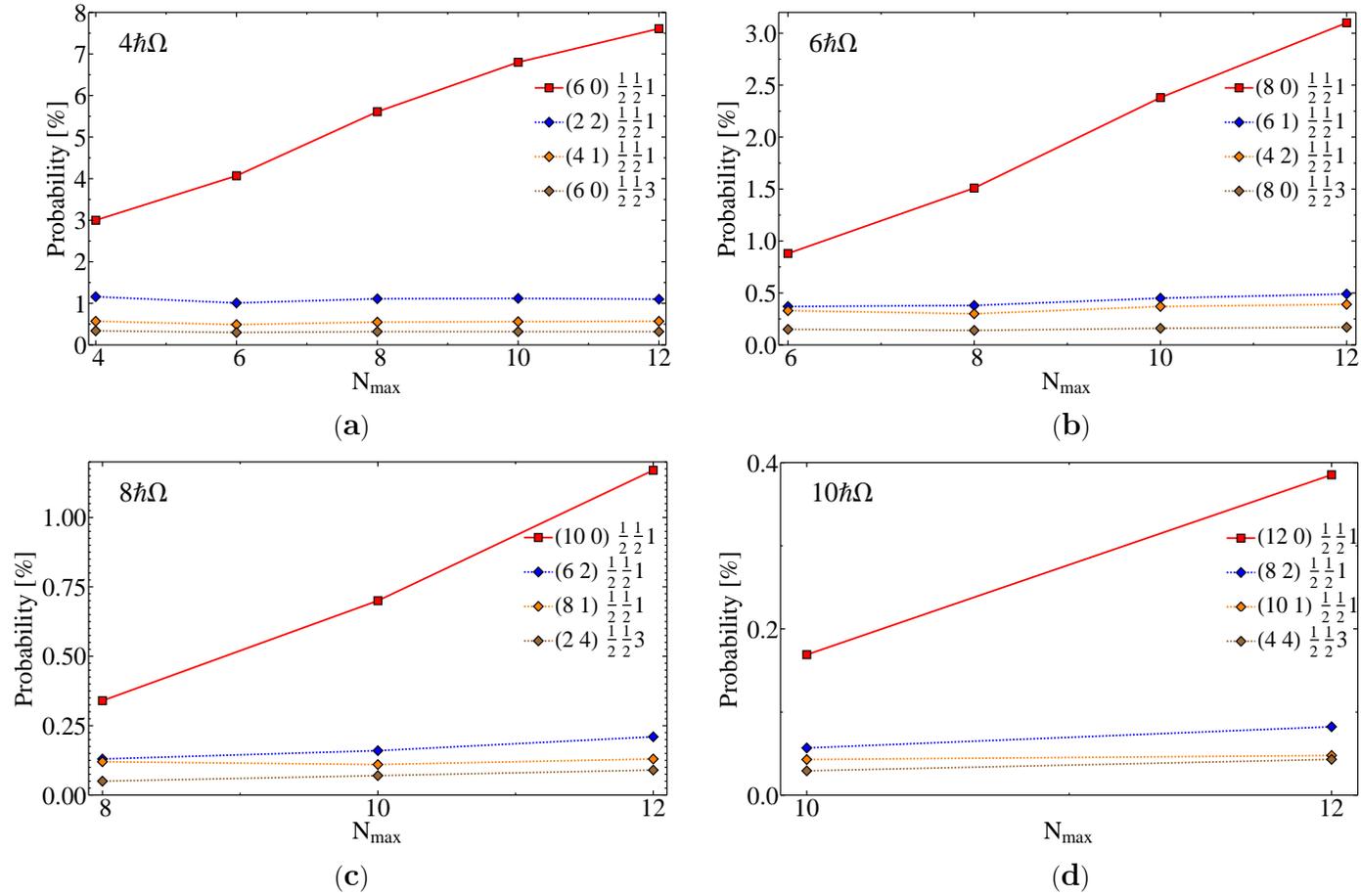


Figure 4: Probabilities of the most important $(\lambda\mu) (S_p S_n S)$ components in ${}^6\text{Li}$ at $4\hbar\Omega$ subspace (a), $6\hbar\Omega$ subspace (b), $8\hbar\Omega$ subspace (c), and $10\hbar\Omega$ subspace as a function of the model space cutoff N_{max} .

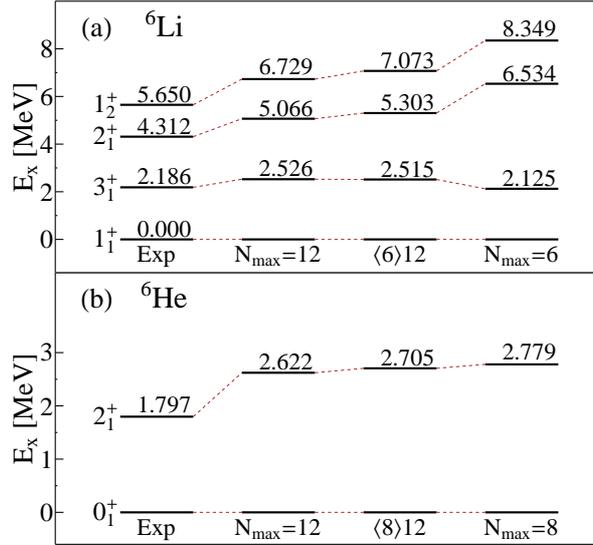


Figure 5: Experimental and theoretical excitation energies: (a) $T = 0$ states of ${}^6\text{Li}$, and (b) the two lowest-lying states of the halo ${}^6\text{He}$ nucleus. Experimental results [24] are given in the first column. The theoretical results shown are for JISP16 and $\hbar\Omega = 20$ MeV in the complete $N_{\max} = 12$ space (second column), symmetry-guided truncated model space (third column) and the complete $N_{\max} = 6$ or 8 spaces (last column). Note the relatively large change in the calculated excitation spectrum of ${}^6\text{Li}$ when N_{\max} is increased from 6 to 12, and that the $\langle 6 \rangle 12$ SA-NCSM results (third column) track the latter closely.

sent from 98% up to 98.7% of the complete-space binding energy in the case of ${}^6\text{Li}$, and reach over 99% for ${}^6\text{He}$. Furthermore, the excitation energies differ only by 11 keV to a few hundred keV from the corresponding complete-space results, see Fig. 5, and the agreement with known experimental data is reasonable over a broad range of $\hbar\Omega$ values.

As illustrated in Table 2, the magnetic dipole moments for ${}^6\text{Li}$ agree to within 0.3% for odd- J values, and 5% for $\mu(2_1^+)$. Qualitatively similar agreement is achieved for $\mu(2_1^+)$ of ${}^6\text{He}$, as shown in Table 3. The results suggest that it may suffice to include all low-lying $\hbar\Omega$ states up to a fixed limit, e. g., $N_{\max}^\perp = 6$ for ${}^6\text{Li}$ and $N_{\max}^\perp = 8$ for ${}^6\text{He}$, to account for the most important correlations that contribute to the magnetic dipole moment.

To explore how close one comes to reproducing the important long-range correlations of the complete $N_{\max} = 12$ space in terms of nuclear collective excitations within

Table 2: Magnetic dipole moments μ [μ_N] and point-particle rms matter radii r_m [fm] of $T = 0$ states of ${}^6\text{Li}$ calculated in the complete $N_{\max} = 12$ space and the $\langle 6 \rangle 12$ subspace for JISP16 and $\hbar\Omega = 20$ MeV. The experimental value for the 1^+ ground state is known to be $\mu = +0.822 \mu_N$ [24].

	1_1^+	3_1^+	2_1^+	1_2^+
μ				
Full $N_{\max} = 12$	0.838	1.866	0.960	0.336
$\langle 6 \rangle 12$	0.840	1.866	1.015	0.337
rms				
Full $N_{\max} = 12$	2.146	2.092	2.257	2.373
$\langle 6 \rangle 12$	2.139	2.079	2.236	2.355

Table 3: Selected observables for the two lowest-lying states of ${}^6\text{He}$ obtained in the complete $N_{\text{max}} = 12$ space and $\langle 8 \rangle 12$ model subspace for JISP16 and $\hbar\Omega = 20$ MeV.

	$N_{\text{max}} = 12$	$\langle 8 \rangle 12$
$B(E2; 2_1^+ \rightarrow 0_1^+) [e^2\text{fm}^4]$	0.181	0.184
$Q(2_1^+) [e\cdot\text{fm}^2]$	-0.690	-0.711
$\mu(2_1^+) [\mu_N]$	-0.873	-0.817
$r_m(2_1^+) [\text{fm}]$	2.153	2.141
$r_m(0_1^+) [\text{fm}]$	2.113	2.110

the symmetry-truncated spaces under consideration, we compared observables that are sensitive to the tails of the wavefunctions; specifically, the point-particle rms matter radii, the electric quadrupole moments and the reduced electromagnetic $B(E2)$ transition strengths that, in addition, could hint at rotational features [25]. As Table 3 clearly shows, the complete-space results for these observables are remarkably well reproduced by the SA-NCSM for ${}^6\text{He}$ in the restricted $\langle 8 \rangle 12$ space. Similarly, the $\langle 6 \rangle 12$ eigensolutions for ${}^6\text{Li}$ yield results for $B(E2)$ strengths and quadrupole moments that track very closely with their complete $N_{\text{max}} = 12$ space counterparts for all values of $\hbar\Omega$ (Fig. 6). The $B(E2)$ strengths almost double upon increasing the model space from $N_{\text{max}} = 6$ to $N_{\text{max}} = 12$. This result suggests that further expansion of the model space will be needed to reach convergence [26]. The close correlation between the $N_{\text{max}} = 12$ and $\langle 6 \rangle 12$ results is nevertheless impressive. In addition to being in agreement, the results reproduce the challenging sign and magnitude of the ground-state quadrupole moment that is measured to be $Q(1^+) = -0.0818(17) e\cdot\text{fm}^2$ [24].

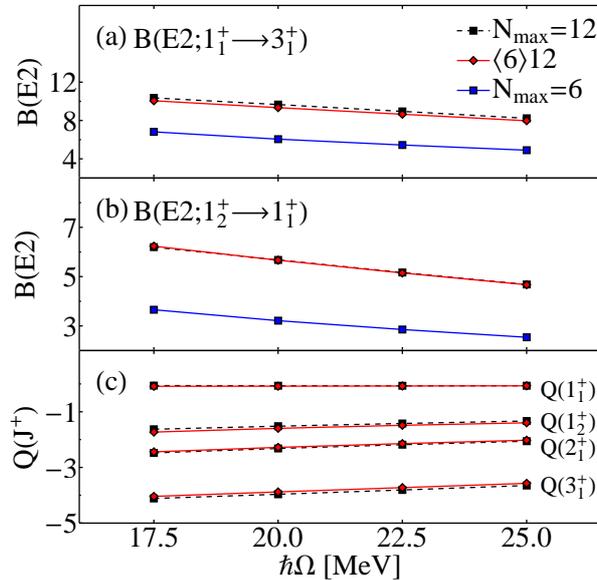


Figure 6: Electric quadrupole transition probabilities in units of $e^2\text{fm}^4$ [(a) and (b), as shown], and quadrupole moments in units of $e\cdot\text{fm}^2$ (c) as a function of $\hbar\Omega$ for $T = 0$ states of ${}^6\text{Li}$ calculated using JISP16 in the complete $N_{\text{max}} = 12$ space (dashed black line), the complete $N_{\text{max}} = 6$ space (solid blue line), and symmetry-truncated $\langle 6 \rangle 12$ (solid red line) model spaces. Note that while the $N_{\text{max}} = 6$ results differ considerably from their $N_{\text{max}} = 12$ counterparts, in all cases the latter are nearly indistinguishable from the truncated $\langle 6 \rangle 12$ results. Experimentally, $B(E2; 1_1^+ \rightarrow 3_1^+) = 25.6(20) e^2\text{fm}^4$ [24].

Finally, the results for the rms matter radii of ${}^6\text{Li}$, listed in Table 2, agree to within 1%.

The differences between truncated-space and complete-space results are found to be essentially insensitive to the choice of $\hbar\Omega$ and appear sufficiently small as to be inconsequential relative to the residual dependences on $\hbar\Omega$ and on N_{max} (see Fig. 6). Since the NN interaction dominates contributions from three-nucleon forces (3NFs) in light nuclei, except for selected cases [27–29], we expect our results to be robust and carry forward to planned applications that will include 3NFs.

5 Conclusion

We have developed a novel approach that capitalizes on advances being made in *ab initio* methods while exploiting exact and partial symmetries of nuclear many-body system. Using this approach we have demonstrated that the low-lying eigenstates of ${}^6\text{Li}$, ${}^8\text{Be}$, ${}^{12}\text{C}$, and ${}^{16}\text{O}$, which were obtained using the JISP16 and $N^3\text{LO}$ NN interaction, exhibit a strong dominance of few intrinsic spin components and carry an intriguingly simple pattern of dominant deformations. The results very clearly underscore the significance of the $\text{SU}(3)$ scheme, LS -coupling, and underlying symplectic symmetry in enabling an extension, through symmetry-guided model space reductions, of *ab initio* methods to heavier nuclei beyond ${}^{16}\text{O}$.

Acknowledgments

This work was supported in part by the US NSF (OCI-0904874, OCI-0904782, PHY-0904782), the US Department of Energy (DE-SC0005248, DE-SC0008485, DE-FG02-87ER40371), and the Southeastern Universities Research Association. This research used computing resources of the Louisiana Optical Network Initiative, LSU's Center for Computation & Technology, and the National Energy Research Scientific Computing Center, which is supported by the Office of Science of the US Department of Energy under Contract No. DE-AC02-05CH11231.

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