



Chiral and deconfining phase transitions from holographic QCD study

Zhen Fang^{a,b,*}, Song He^{c,d,a}, Danning Li^a

^a Key Laboratory of Theoretical Physics, Institute of Theoretical Physics, Chinese Academy of Science, Beijing 100190, PR China

^b University of Chinese Academy of Sciences, Beijing, PR China

^c Max Planck Institute for Gravitational Physics (Albert Einstein Institute), Am Mühlenberg 1, 14476 Golm, Germany

^d Yukawa Institute for Theoretical Physics, Kyoto University, Kitashirakawa Oiwakecho, Sakyo-ku, Kyoto 606-8502, Japan

Received 12 February 2016; received in revised form 25 March 2016; accepted 2 April 2016

Available online 7 April 2016

Editor: Hong-Jian He

Abstract

A preliminary quantitative study to match the lattice QCD simulation on the chiral and deconfining phase transitions of QCD in the bottom-up holographic framework is given. We constrain the relation between dilaton field ϕ and metric warp factor A_e and get several reasonable models in the Einstein-Dilaton system. Using the potential reconstruction approach, we solve the corresponding gravity background. Then we fit the background-related parameters by comparing the equation of state with the two-flavor lattice QCD results. After that we study the temperature dependent behavior of Polyakov loop and chiral condensate under those background solutions. We find that the results are in good agreement with the two-flavor lattice results. All the studies about the equation of state, the Polyakov loop and the chiral condensate signal crossover behavior of the phase transitions, which are consistent with the current understanding on the QCD phase transitions with physical quark mass. Furthermore, the extracted transition temperatures are comparable with the two-flavor lattice QCD results.

© 2016 The Authors. Published by Elsevier B.V. This is an open access article under the CC BY license (<http://creativecommons.org/licenses/by/4.0/>). Funded by SCOAP³.

* Corresponding author at: Key Laboratory of Theoretical Physics, Institute of Theoretical Physics, Chinese Academy of Science, Beijing 100190, PR China.

E-mail addresses: fangzhen@itp.ac.cn (Z. Fang), hesong17@gmail.com (S. He), lidn@itp.ac.cn (D. Li).

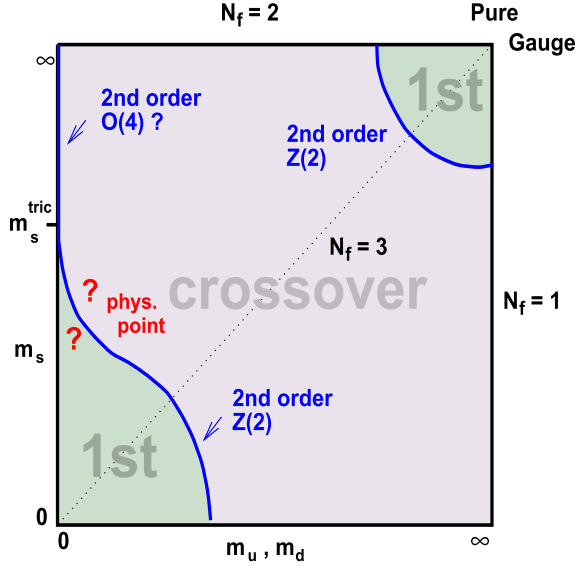


Fig. 1. The expected phase diagram in the current quark mass space with degenerate u, d quark masses (taken from [2]).

1. Introduction

Spontaneous chiral symmetry breaking and color confinement are the two most important properties of the vacuum of Quantum Chromodynamics (QCD), which is widely accepted as the fundamental theory of the strong interaction. At sufficiently high temperature and/or density, it is believed that phase transition might happen in the system, including the restoration of chiral symmetry and the release of color degrees of freedom. At present, to understand the phase structure of these two phase transitions is attracting more and more attention in both non-perturbative QCD study and cosmology [1].

Generally, the properties of the two phase transitions would depend sensitively on the intrinsic quantities of the system. For example, the chiral phase transition is well defined as a true phase transition only in the chiral limit, i.e., zero quark mass limit, while the deconfining phase transition should be in a totally opposite limit, i.e., the infinite quark mass limit. This is because only in these two limits the chiral symmetry and Z_3 center symmetry, the breaking and restoration of which are related to the phase transitions, become the exact symmetries of QCD. In these limits, chiral condensate $\langle \bar{\psi}\psi \rangle$ and Polyakov loop $\langle L \rangle$ could be well defined as order parameters for chiral and deconfining phase transition respectively. In the physical quark mass region, there are no exact symmetries and the phase transition might turn to a rapid but continuous crossover transition.

Based on theoretical consideration and lattice QCD simulation [2–4], a possible $2 + 1$ flavor phase diagram in the current quark mass plane is summarized in the sketch (sometimes called “Colombia Plot”) shown in Fig. 1 [2]. In this plot, there are two regions with first-order phase transition, i.e., near the chiral limit region ($m_u = m_d = m_s \simeq 0$) and near the infinite quark mass region ($m_u = m_d = m_s \simeq \infty$). In the intermediate region it is expected to be a crossover transition. There are two second-order lines as the boundaries between the first-order regions and the crossover region. It is noted that in the region of two light flavors where $m_u = m_d \simeq O(\text{MeV})$

and $m_s = \infty$, even very small quark mass would drive the second-order transition in the chiral limit to a crossover transition with finite quark mass (in analogy to the $O(4)$ σ model [5] noting that $SU(2)_L \times SU(2)_R \simeq O(4)$). In this paper, as a preliminary try, we will focus on the behavior of phase transitions in this area.

In addition to the quark mass and flavor dependent behavior of QCD phase transition, one of the most important things for both experimental investigations of quark gluon plasma and theoretical studies of thermal QCD is to estimate the transition temperature T_c . As the improvement of lattice computation in recent years, most of the results about the transition temperatures converge towards $T_c \simeq 145 \sim 165$ MeV in 2+1 flavor QCD [4]. In the two-flavor case, the transition temperature is around 170 MeV when extrapolated to the chiral limit, while in the three-flavor case it is about 155 MeV [2]. The transition temperature would increase with increasing quark mass. Another interesting topic is about the relation of chiral and deconfining transitions. In this aspect, discrepancies exist for different theoretical studies. Besides lattice simulations (see references in [4]), there are also other efforts such as using effective models or by functional methods [6–9]. In this work, we will try to use the holographic approach to study the thermal phase transition of QCD and try to provide more understanding on it.

In recent decades, the discovery of the anti-de Sitter/conformal field theory (AdS/CFT) correspondence and the conjecture of the gravity/gauge duality [10–12] shed new light on the strong coupling problem of gauge theory. Based on this idea, people have extended it to study gauge theory like QCD, in which conformal symmetry is broken dynamically at low energy. By breaking the conformal symmetry in different ways, many efforts have been made towards more realistic holographic description of the low energy phenomena of QCD, such as in hadron physics [13–29] and hot/dense QCD [30–46], both in top-down approaches and in bottom-up approaches (see [47–51] for reviews).

For QCD phase transitions, most of the bottom-up studies [52–78] focus only on deconfining phase transition. To add chiral aspects, the soft-wall model [79] provides a starting point, since the model itself and its extended ones [22,24–29] have been shown to characterize the hadron spectra and other quantities quite well. Furthermore, in these models the chiral condensate, which is the order parameter of chiral phase transition, was introduced to realize the spontaneous chiral symmetry breaking of QCD vacuum. However, unlike the Nambu–Jona–Lasinio (NJL) model [80,81], the value of the chiral condensate was often taken as a free parameter to fit the hadron spectra at zero temperature instead of being self-consistently solved from the model itself. Noting that the IR boundary condition may require the dependence of chiral condensate on quark mass, the authors of [82] extended the soft-wall model to the finite temperature case and solved the temperature dependent chiral condensate self-consistently. In the previous work [83,84], we extended their method and tried to get more constraints on the model from chiral aspects of QCD phase transition. There we showed how to get the correct mass dependent behavior of chiral phase transition as shown in Fig. 1. Furthermore, we also see that under the AdS–Schwarzschild (AdS–SW) black hole background new constraints on the dilaton field come up: it should be negative at certain UV scale in addition to the IR constraints from meson spectra. However, since in that work we only focused on the chiral phase transition and used the AdS–SW black hole solution as the background geometry, we did not consider the deconfining phase transition there. In this work, we will try to grasp the main requirement in describing chiral and deconfining phase transition simultaneously, and a holographic QCD (hQCD) model will be built up to characterize the behavior of the two aspects of phase transitions.

The paper is organized as follows. In Sec. 2, we describe the Einstein-dilaton system, which has been used to characterize the deconfining phase transition in previous studies. We try to use

the potential reconstruction method to construct several models. Then we study thermodynamics of these models and compare the results with two-flavor lattice QCD simulations in Sec. 3. After fixing the parameters by the description of equation of state, we study the temperature dependent behavior of Polyakov loop and chiral condensate, and also compare them with the results from lattice in Sec. 4. In Sec. 5, we give a short discussion and conclusion.

2. Gravity setup

As mentioned above, in bottom-up holographic studies, the deconfining phase transition has already been widely investigated in the Einstein-dilaton system [60–76]. In addition, in [83,84] we showed the possibility to characterize chiral symmetry breaking and its restoration in soft-wall model. In light of these researches, we expect that by combining these two systems it should be possible to describe the two most important aspects of QCD phase transition simultaneously. Therefore, we consider the following action in string frame

$$S = S_b + S_m, \quad (1)$$

$$S_b = \frac{1}{16\pi G_5} \int d^5x \sqrt{-g^S} e^{-2\phi} \left[R^S + 4\partial_\mu \phi \partial^\mu \phi - V_S(\phi) \right], \quad (2)$$

$$S_m = - \int d^5x \sqrt{-g^S} e^{-\phi} \text{Tr} \left[\nabla_\mu X^\dagger \nabla^\mu X + V_X(|X|) \right], \quad (3)$$

where S_b is the background sector and S_m the matter sector. The index S in the integrand denotes the string frame and G_5 is the 5D Newton constant. There are two scalar fields, i.e., the dilaton field ϕ and the bulk scalar field X which is dual to the $\bar{q}q$ condensate of QCD. V_S and V_X represent the dilaton potential and the bulk scalar potential respectively. Here the leading term of V_X is the mass term $m_5^2 X X^\dagger$ and the bulk scalar mass $m_5^2 = -3$ can be determined from the AdS/CFT prescription $m_5^2 L^2 = (\Delta - p)(\Delta + p - 4)$ by taking $\Delta = 3$, $p = 0$ [12].

In [70,71], we showed that the Einstein-dilaton system can describe pure gluon thermodynamics quite well. After adding the flavor sector S_m , we found that the above system can describe the meson spectra which are consistent with experimental data [28,29]. However, extending this model to the finite temperature case and trying to solve this gravity-two-scalar coupled system is quite a complicated work.¹ Actually, in [60,61,67], the authors also tried to study thermodynamics of QCD with flavors in the Einstein-dilaton system. Following this logic and as a preliminary try, we will solve the Einstein-dilaton action S_b as the geometrical background and take the matter action S_m as a probe, which should be considered as an approximation of the full system. As pointed out above, we consider chiral and deconfining phase transitions in the two-flavor case ($N_f = 2$), so the background geometry will be constrained by $N_f = 2$ thermodynamics before studying the temperature dependent behavior of chiral condensate and Polyakov loop. In this section, we will first outline the necessary framework of the Einstein-dilaton system, and then try to constrain the background and give several models for study.

2.1. Equation of motion for background geometry

Given the background action S_b as shown in Eq. (2), if one knows the dilaton potential $V_S(\phi)$, then the whole system could be solved numerically. This is the usual approach to deal with this

¹ The authors of [85,86] has done full analysis in a different system other than the soft-wall AdS/QCD model, which has been proved to be successful in describing chiral symmetry breaking and linear confinement of QCD.

system, as can be seen in [60–69], where by tuning the dilaton potential carefully the authors can describe the QCD equation of state quite well. While in [45,46,70–72], we used an approximate approach, which is usually called potential reconstruction approach, to construct the geometrical background (see also [73,74,87,88]). In this approach, once fixing the dilaton profile [45,46], the metric warp factor [70–74] or the relations between them, the dilaton potential could be solved from the equations of motion, which nevertheless entails a temperature dependence of the potential. However, in the region concerned, the temperature dependence of dilaton potential is very weak, so it can be seen as an approximation of the potential fixing approach (see also the discussion in [89]). Furthermore, it turns out to be easier to generate the background solution. Hence, in this work, we will use this approach and try to get the necessary ingredients in describing QCD phase transitions holographically. Here, we first review how to use the potential reconstruction approach to obtain solutions in the 5D Einstein-dilaton system given in Eqs. (1)–(3).

For convenience, we give the relevant formulas in Einstein frame. By conformal transformation of the metric

$$g_{\mu\nu}^S = e^{\frac{4\phi}{3}} g_{\mu\nu}^E, \quad (4)$$

the action in Einstein frame is derived as

$$S_b = \frac{1}{16\pi G_5} \int d^5x \sqrt{-g^E} \left[R^E - \frac{4}{3} \partial_\mu \phi \partial^\mu \phi - V_E(\phi) \right], \quad (5)$$

with

$$V_E(\phi) = e^{\frac{4\phi}{3}} V_S(\phi). \quad (6)$$

For finite temperature solution, the metric ansatz in string frame and Einstein frame will be taken as follows

$$ds_S^2 = \frac{L^2 e^{2A_s}}{z^2} \left(-f(z) dt^2 + dx^i dx^i + \frac{dz^2}{f(z)} \right), \quad (7)$$

$$ds_E^2 = \frac{L^2 e^{2A_e}}{z^2} \left(-f(z) dt^2 + dx^i dx^i + \frac{dz^2}{f(z)} \right), \quad (8)$$

where L is the radius of AdS_5 , and the relation of the metric warp factor in different frames is $A_s = A_e + 2\phi/3$.

From the action (5), the general Einstein equation can be derived as

$$E_{\mu\nu} + \frac{1}{2} g_{\mu\nu}^E \left(\frac{4}{3} \partial_\mu \phi \partial^\mu \phi + V_E(\phi) \right) - \frac{4}{3} \partial_\mu \phi \partial_\nu \phi = 0, \quad (9)$$

from which we obtain the non-zero components:

$$A_e''(z) + A_e'(z) \left(\frac{f'(z)}{2f(z)} - \frac{2}{z} + A_e'(z) \right) - \frac{f'(z)}{2zf(z)} + \frac{2}{z^2} + \frac{2}{9} \phi'(z)^2 + \frac{L^2 e^{2A_e(z)} V_E(\phi(z))}{6z^2 f(z)} = 0, \quad (10)$$

$$-\frac{9f'(z)A_e'(z)}{4f(z)} - 9A_e'(z)^2 + \frac{18A_e'(z)}{z} + \frac{9f'(z)}{4zf(z)} - \frac{9}{z^2} + \phi'(z)^2 - \frac{3L^2 e^{2A_e(z)} V_E(\phi(z))}{4z^2 f(z)} = 0, \quad (11)$$

$$f''(z) + f(z) \left(6A'_e(z)^2 - \frac{12A'_e(z)}{z} + 6A''_e(z) + \frac{12}{z^2} + \frac{4}{3}\phi'(z)^2 \right) + f'(z) \left(6A'_e(z) - \frac{6}{z} \right) + \frac{L^2 e^{2A_e(z)} V_E(\phi(z))}{z^2} = 0. \quad (12)$$

Note that we only need two of the above three equations in which $V_E(\phi)$ is considered as a derived quantity in the potential reconstruction approach. The other one can be used as a consistent check for solutions of the equations of motion. For simplicity, we recombine Eqs. (10)–(12) and obtain the following two simplified equations:

$$f''(z) + 3f'(z) \left(A'_e(z) - \frac{1}{z} \right) = 0, \quad (13)$$

$$A''_e(z) + A'_e(z) \left(\frac{2}{z} - A'_e(z) \right) + \frac{4}{9}\phi'(z)^2 = 0. \quad (14)$$

Also from the action (5), we obtain the dilaton field equation:

$$\phi''(z) + \phi'(z) \left(3A'_e(z) + \frac{f'(z)}{f(z)} - \frac{3}{z} \right) - \frac{3L^2 e^{2A_e(z)} \partial_\phi V_E(\phi(z))}{8z^2 f(z)} = 0. \quad (15)$$

Note that only A_e and ϕ appear in Eq. (14). If one of these two quantities or the relation between them were given, they could be solved from Eq. (14), then f and $V_E(\phi)$ would be obtained by solving Eq. (13) and Eq. (15). Usually only the integral constant in f would appear in the final expression of $V_E(\phi)$. As the integral constant is related to the black hole temperature, this indicates that $V_E(\phi)$ would depend on temperature. However, as we can see later, the temperature dependence of $V_E(\phi)$ is very weak, which makes it possible to consider this approach as an approximation of fixing dilaton potential. Taking specific profile of ϕ and A_e , we studied thermodynamics of the pure gluon system in [45,46,70,71] and found that the results are in good agreement with the quenched lattice QCD results. In this work, we aim at characterizing thermodynamics and phase transitions in $N_f = 2$ QCD with finite quark mass, which shows a crossover transition instead of a first-order one such as appears in the pure gluon system. Thus, we will attempt to acquire new constraints on A_e , ϕ and construct the gravity background for characterizing two-flavor QCD thermodynamics.

2.2. UV and IR constraints on background geometry and dilaton

To tackle the gravity background, ϕ and A_e (or equivalently A_s) should be specified as the input of the background Eqs. (13)–(15). Starting from a specific relation, an asymptotic AdS black hole solution will be obtained. In this section, we will first find the UV and IR constraints of the corresponding quantities, and then try to build reasonable gravity background models. In our convention, the UV and IR region are corresponding to $z \sim 0$ and large z respectively.

Firstly, as noted in [28,29,70,71], the Einstein-dilaton system should be closely related to the gluon dynamics, which means that the dimension of the dilaton field should be $\Delta = 2$ or $\Delta = 4$, which is equivalent to requiring the leading UV asymptotic behavior of ϕ to be z^2 or z^4 forms according to the holographic dictionary.

Secondly, as shown in [61,65,70,71,86], the UV asymptotic AdS region is related to the high temperature behavior of the thermodynamic system. The asymptotic AdS property guarantees the system to approach conformal gas at very high temperature. However, in order to have a correct description of the low temperature behavior, the IR asymptotic behavior of the background fields

should be carefully tuned. From the previous studies [28,29,61,65,70,71,86], we see that if the IR behavior of A_s or ϕ is of the quadratic form z^2 , then the black hole solutions would have a minimal temperature and the system shows a first-order phase transition. However, if the flavor sector and quark mass are taken into consideration, the QCD phase transition will turn into a crossover one without a real phase transition. Therefore, we need a gravity background which can link the high temperature phase with the low temperature phase. In this case, we tune the IR behavior of the fields carefully and find that when they approach a constant at IR the temperature corresponding to the AdS black hole solution can go down to zero continuously, as will be shown in Fig. 2.²

Thirdly, the main motivation of this paper is to study deconfining phase transition together with chiral phase transition. In our previous study [83,84], we found that new constraints on dilaton field could be obtained from the chiral aspects of QCD phase transition. As here the gravity background is no longer pure AdS, one needs another similar constraint on $5A_s - \phi$, noting that $e^{5A_s - \phi}$ comes from the sector $\sqrt{-g}S e^{-\phi} = \frac{1}{5}e^{5A_s - \phi}$ of the matter action (3) which will be considered in Sec. 4.2. In [83,84] we showed that a negative part of dilaton at certain scale not far from UV is necessary to obtain the correct chiral phase transition behavior in the pure AdS background. Accordingly, here we only require $5A_s - \phi$ to be positive for simplicity.

As a short summary, we need a UV leading z^2 or z^4 configuration of $5A_s - \phi$ or ϕ itself, and we will simply set the IR behavior of them to approach a positive constant for a possible crossover transition. The simplest choice to interpolate the UV and IR behavior required above is of $\tanh$ form, and the next order term will be retained to fit the correct intermediate behavior of thermal transition. Hence, we give the following four possible ansatz, and set them to be model A1 (A2) and model B1 (B2),

$$5A_s(z) - \phi(z) = \beta \tanh(\mu^2 z^2 + \nu^4 z^4) \quad (\text{model A1}), \quad (16)$$

$$\phi(z) = \beta \tanh(\mu^2 z^2 + \nu^4 z^4) \quad (\text{model A2}), \quad (17)$$

$$5A_s(z) - \phi(z) = \beta \tanh(\mu^4 z^4 + \nu^6 z^6) \quad (\text{model B1}), \quad (18)$$

$$\phi(z) = \beta \tanh(\mu^4 z^4 + \nu^6 z^6) \quad (\text{model B2}). \quad (19)$$

Here β, μ, ν are model parameters and will be fixed later by comparing the results of equation of state with those from lattice simulations.

It is easy to obtain the UV asymptotic behavior of $\phi(z)$ under the above settings as follows

$$\phi(z \rightarrow 0) \sim \frac{3\beta\mu^2 z^2}{7} + \dots \quad (\text{model A1}), \quad (20)$$

$$\phi(z \rightarrow 0) \sim \beta\mu^2 z^2 + \dots \quad (\text{model A2}), \quad (21)$$

$$\phi(z \rightarrow 0) \sim \frac{3\beta\mu^4 z^4}{7} + \dots \quad (\text{model B1}), \quad (22)$$

$$\phi(z \rightarrow 0) \sim \beta\mu^4 z^4 + \dots \quad (\text{model B2}). \quad (23)$$

² Here we note that this requirement is in contradiction with constraints from meson spectrum, where an IR quadratic dilaton is needed to produce the Regge behavior of meson spectra. However, in this manuscript, we just want to grasp the most important ingredients in describing QCD phase transitions and not to lay emphasis on the mass spectrum. Furthermore, we note that in [90], by adding an extra scalar field, it is possible to get pure AdS together with IR quadratic dilaton field which is consistent with the spectrum calculation. One might combine these two aspects in one model through this way. We will leave the more careful study to the future.

Table 1

Value of parameters in model A1 (A2) and model B1 (B2).

	β	μ (GeV)	ν (GeV)	G_5
Model A1	2.5	0.5	0.46	1.06
Model A2	1.8	0.4	0.42	1.06
Model B1	1.8	0.59	0.52	1.1
Model B2	1.3	0.55	0.38	1.1

From these results, we see that the conformal dimension of the dilaton field is $\Delta = 2$ in A1 (A2) model and $\Delta = 4$ in B1 (B2) model and they all satisfy the BF bound. The dilaton field with $\Delta = 4$ can be dual to the gauge invariant dimension-4 gluon condensate, while the dilaton with $\Delta = 2$ does not correspond to any local, gauge invariant operator in QCD. Although there have been many discussions in recent years of the possible relevance of a dimension-two condensate in the form of a gluon mass term [91–93], it is still not clear whether we can associate ϕ with a dimension-2 operator, because the AdS/CFT correspondence requires that the bulk fields should be dual to gauge-invariant local operators. Candidates like non-local gauge-invariant operators $F_{\mu\nu}(D^2)^{-1}F^{\mu\nu}$ [94] are proposed in [20]. To include such kind of possibilities, here we also consider the $\Delta = 2$ case. Despite of all these issues, we will just go on to find the physical implications on the thermal QCD phase transition from the holographic studies.

3. Equation of state for the hQCD model

After giving the main framework of the hQCD model, in this section, we will investigate the equation of state in these models given in Eqs. (16)–(19). There are four parameters β , μ , ν , G_5 in these models. Once these model parameters are given, one can solve the metric and dilaton background from the equation of motion. Once the background solutions are solved, roughly, there are two ways to extract the thermodynamical quantities, such as entropy density, sound speed, pressure and so on. One is to use the equivalence of partition function and extract the thermodynamical quantities from the on-shell action (for details, see [66]). The other one is to integrate the other thermodynamical quantities from the entropy density, which could be easily evaluated using the famous Bekenstein–Hawking formula. In general, in the former approach, one has to introduce counter terms to cancel the UV divergence, which might introduce scheme dependence to the results. To cancel the counter term dependence, only the difference of the quantities in different phases are physically meaningful. Since in [95] the authors showed that these two approaches should give the same results, to simplify the calculation, we will follow the latter approach as in [60,61]. The details to extract the equation of states will be given later. We fix the parameters β , μ , ν and G_5 in the models by comparing the calculated results with the $N_f = 2$ lattice ones from [96], and list them in Table 1 for later use.

3.1. Black hole solutions and associated thermodynamics

Before going to the details, we will list general formulas of some thermodynamic quantities. Here we are interested in a series of solutions whose UV behavior is asymptotic AdS₅. We also impose the requirements: $f(0) = 1$, and $\phi(z)$, $f(z)$, $A_s(z)$ are regular from $z = 0$ to $z = z_h$. Here z_h is the black hole horizon with $f(z_h) = 0$.

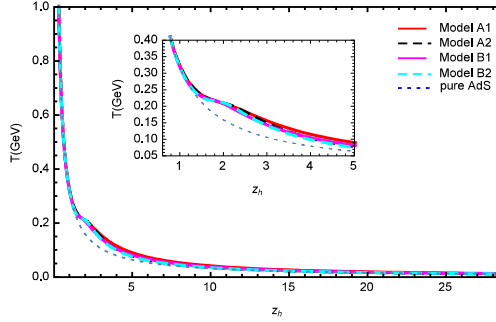


Fig. 2. The temperature as a function of horizon z_h in model A1 (A2) and model B1 (B2).

The Hawking temperature of the black hole solution is defined by

$$T = \frac{|f'(z_h)|}{4\pi}. \quad (24)$$

The background Eq. (13) can be simplified by defining $f(z)$ as

$$f(z) = 1 - \frac{h(z)}{h(z_h)} \quad (25)$$

with $h(0) = 0$. Then the temperature formula (24) becomes

$$T = \frac{h'(z_h)}{4\pi h(z_h)}. \quad (26)$$

Substituting Eq. (25) in Eq. (13), we obtain a simplified equation of $h(z)$:

$$h'(z) - e^{-3A_e(z)} z^3 = 0. \quad (27)$$

Now we are ready to solve the background system completely. In Fig. 2, we present the temperature as a function of the horizon z_h in model A and B. We see that the temperature decrease monotonously with z_h from high ones to zero. The fact that the behavior of $T(z_h)$ is almost same in all the four models quantitatively means that the black hole temperature is not sensitive to the relations of A_s and ϕ we have chosen. That is critical for a consistent realization of the crossover transition and also important as an indication that the model is robust enough in characterizing the transition features of thermal QCD. Compared with the pure AdS–SW black hole case, there is a small z_h region in which the temperature decrease slightly slowly in these models, as can be seen from the insert of Fig. 2.

The dilaton potential can be derived from Eqs. (10)–(12), and the results in model A1 and B1 are shown in Fig. 3 with several different temperatures. From Fig. 3 one can see that the dilaton potential will approach to the negative cosmological constant when ϕ goes to zero, which is consistent with the UV asymptotic AdS boundary. In the IR region, the potential will be dependent on the temperature due to the potential construction. However, since the relevant physical region is from the boundary $z = 0$ to the horizon $z = z_h$, we only plot the section from $\phi = 0$ to $\phi = \phi(z_h)$. One can see that in this region the dilaton potential almost does not change with temperature. In this sense, it might be possible to build up a model with fixing dilaton potential which could produce similar results as that in this work.

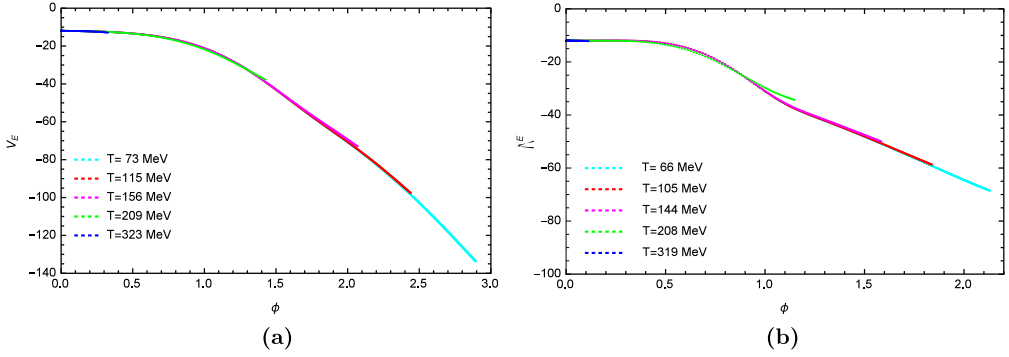
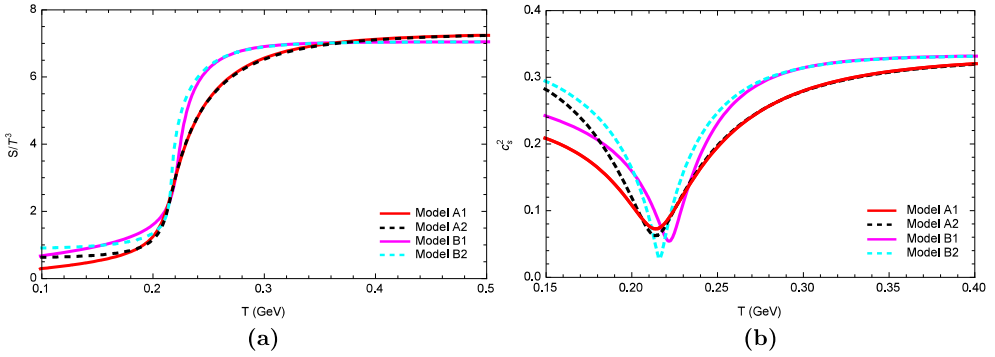


Fig. 3. The dilaton potential in model A1 (a) and model B1 (b).

Fig. 4. Temperature dependence of the scaled entropy density s/T^3 (a) and the square of sound speed c_s^2 (b) in model A1 (A2) and model B1 (B2).

3.2. Entropy density and sound speed

Based on the classical Bekenstein–Hawking formula and using the metric ansatz (8) in Einstein frame, the formula of the black-hole entropy density is derived as

$$s = \frac{A_{\text{area}}}{4G_5 V_3} \Big|_{z_h} = \frac{L^3}{4G_5} \frac{e^{3A_e}}{z^3} \Big|_{z_h} \quad (28)$$

with A_{area} the area of horizon, G_5 the 5D Newton constant and V_3 the space volume. The sound speed is defined as

$$c_s^2 = \frac{d \ln T}{d \ln s}. \quad (29)$$

Given the formula of temperature (24) and entropy density (28), the speed of sound can be obtained. It has been well known that for conformal system, $c_s^2 = 1/3$, while for non-conformal system, c_s^2 will deviate from $1/3$. From Eq. (29), one can see that the speed of sound is independent of the normalization of the 5D Newton constant G_5 and the space volume V_3 . Fig. 4 presents the scaled entropy density s/T^3 and the sound speed c_s^2 as a function of T in model A1 (A2) and model B1 (B2).

One can see from Fig. 4 (a) that the crossover transition is obvious, and the scaled entropy density increases with temperature. There is a small temperature region where $\frac{s}{T^3}$ increases rapidly, which indicates a sudden release of degrees of freedom. At high temperature, $\frac{s}{T^3}$ approaches to a finite value $\frac{\pi^3}{4G_5}$ which can be obtained from asymptotic analysis. From thermal QCD one knows that the scaled entropy density at high temperature would approach the Stefan-Boltzmann limit which is related to the degrees of freedom of the system. Thus, in this sense, if we tune G_5 and fix the high temperature value of $\frac{s}{T^3}$ comparable with the $N_f = 2$ lattice QCD result, the flavor effects will partly incorporated. In Fig. 4 (b), we plot the square of sound speed varying with temperature. All these four models show a minimal value within the temperature region $210 \sim 230$ MeV. This also indicates a crossover behavior which is consistent with the entropy density. At high temperature, the square of sound speed goes to the conformal value $1/3$. It should be noted that our high temperature analysis of the thermal quantities may be irrelevant from the point of view of holography because at very high temperatures the property of asymptotic freedom will be attained, which makes the low energy supergravity approximation of the full string theory used in the AdS/CFT (weak/strong) duality invalid. However, from the results we obtained, one finds that the high temperature behavior is consistent with the lattice or other thermal QCD results, at least in a temperature region (up to 0.4 GeV) accessible to the lattice simulations. So we assume that when the temperature is not far below or far above the thermal transition region, the hQCD model can give an approximately good description for the thermal QCD transition.

3.3. Pressure, energy density and trace anomaly

The pressure p is given by the formula

$$\frac{dp}{dT} = s. \quad (30)$$

Numerically, we transform Eq. (30) into $p'(z_h) = s(z_h)T'(z_h)$ which is solved by giving the initial condition $p(z_h = 30 \text{ GeV}^{-1}) = 0$, that is, we set $p = 0$ at $T \simeq 0$. The energy density $\epsilon = -p + sT$ and the trace anomaly $\epsilon - 3p$ are then obtained from the entropy density and the pressure.

We show the numerical results in Fig. 5 and Fig. 6. The lattice results with two light flavors [96] are added in for fitting and lattice results with $2 + 1$ flavors [97] also added for comparison. In Fig. 5, we show the behavior of the scaled pressure and energy density with respect to temperature in the four models. The temperature dependence of the scaled trace anomaly is presented in Fig. 6. The color band denotes the interpolation results of lattice simulations. One can see that the numerical results calculated from the four models are consistent with the $N_f = 2$ lattice results and a similar crossover behavior emerged in a small region of temperature, which is consistent with the behavior of entropy density and sound speed in Fig. 4. The scaled pressure and energy density, which is associated with the degrees of freedom, increase with temperature gradually. At high temperature, the trace anomaly goes to zero, which indicates the system reaches asymptotically to the conformal gas. Qualitatively all the models give consistent results compared with the lattice data, and yet quantitatively the model A1 and A2 fit the lattice results much better than the model B1 and B2.

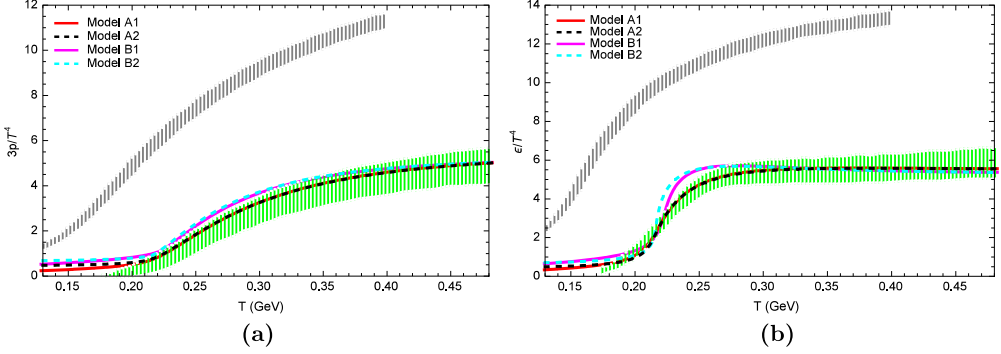


Fig. 5. The scaled pressure $3p/T^4$ and energy density ϵ/T^4 as a function of T in model A1 (A2) and B1 (B2). The green band is the lattice interpolation results with two flavors [96], and the gray band is the lattice results with 2 + 1 flavors [97]. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

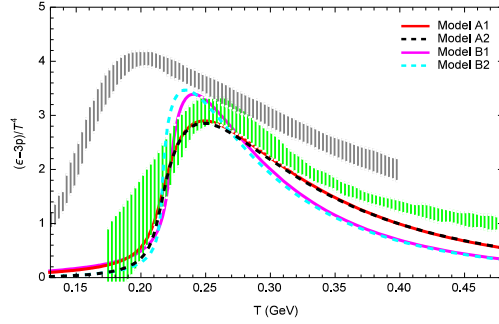


Fig. 6. The scaled trace anomaly $(\epsilon - 3p)/T^4$ as a function of T in model A1 (A2) and B1 (B2). The green band is the lattice interpolation results with two flavors [96], and the gray band is the lattice results with 2 + 1 flavors [97]. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

4. Chiral and deconfining phase transition

In the previous section, we have already seen that the holographic model constructed in Sec. 2 can describe the equation of state in $N_f = 2$ thermal QCD quite well. Qualitatively, these results signal a crossover transition in a similar temperature region. To test the hQCD model further and get more information of the thermal transition, we study the temperature dependent behavior of the order parameters for chiral and deconfining phase transitions in this section.

4.1. Deconfining phase transition from Polyakov loop

The Polyakov loop operator is a useful quantity to investigate the phenomenon of deconfinement. It is defined as

$$L(T) = \frac{1}{N_c} \text{Tr} \text{P} \exp \left(ig \int_0^{1/T} \hat{A}_0(\tau, \vec{x}) d\tau \right) \quad (31)$$

with N_c the color number, P denoting path ordering and the trace computed in the fundamental representation. As an order parameter for the center symmetry of the gauge group, its expectation value $\langle L(T) \rangle$ vanishes in the confined phase, as the center symmetry guarantees, while in the deconfined phase $\langle L(T) \rangle \neq 0$, which implies the center symmetry is breaking. In the holographic framework, the Polyakov loop is related to the action of the worldsheet which wraps the imaginary time circle. Schematically, we can write

$$\langle L(T) \rangle = \int DX e^{-S_w}. \quad (32)$$

In the large N_c and strong coupling limit, the expectation value of the Polyakov loop $\langle L(T) \rangle \sim e^{-S_{\text{NG}}}$ with S_{NG} the Nambu–Goto action for the string worldsheet³:

$$S_{\text{NG}} = \frac{1}{2\pi\alpha_p} \int d^2\eta \sqrt{\det(g_{\mu\nu}^s X_a^\mu X_b^\nu)}, \quad (33)$$

where α_p is the string tension, $g_{\mu\nu}^s$ the metric in the string frame, and X_a^μ the embedding function of the worldsheet in the target space. μ, ν are $5D$ space-time indices and a, b denote the worldsheet coordinates. Using the string frame metric (7), we yield

$$S_{\text{NG}} = \frac{g_p}{\pi T} \int_0^{z_h} dz \frac{e^{2A_s}}{z^2} \sqrt{1 + f(z)(\vec{x}')^2} \quad (34)$$

with $g_p = \frac{L^2}{2\alpha_p}$. The prime indicates the derivative with respect to z . From the action S_{NG} , the equation of motion for $\vec{x}$ is derived as

$$\left[\frac{e^{2A_s}}{z^2} f(z) \vec{x}' / \sqrt{1 + f(z)(\vec{x}')^2} \right]' = 0. \quad (35)$$

By substituting the constant solution of the above equation in the action S_{NG} , the minimal worldsheet is obtained as

$$S_0 = c_p + S'_0 = c_p + \frac{g_p}{\pi T} \int_0^{z_h} dz \left(\frac{e^{2A_s}}{z^2} - \frac{1}{z^2} \right), \quad (36)$$

where c_p is a normalization constant depending on the regularization scheme. Note that the last term of the integrand is a counter term which regularize the UV divergence of the original integral. Now we get the expectation value of the Polyakov loop

$$\langle L(T) \rangle = w e^{-S_0} = e^{C_p - S'_0} \quad (37)$$

with C_p another normalization constant and w a weight coefficient.

Using Eqs. (36)–(37) and the previous results fixed by the equation of state, we can obtain the temperature dependent behavior of Polyakov loop. The parameters C_p, g_p are fixed by two-flavor lattice results from [98]. We present the numerical results of the expectation value of Polyakov loop as a function of T in Fig. 7 (a). The corresponding parameters we used in the plot are listed

³ Here we follow the calculations in [65,67,95,99] and only consider the simplest version of the string action even when the dilaton field is non-trivial. We remark that the coupling of the dilaton to the Ricci scalar on the world-sheet might affect the results of Polyakov loop, which needs a more careful treatment in the future.

Table 2

Polyakov loop related parameters in model A1 (A2) and model B1 (B2).

	C_p	g_p
Model A1	−0.3	0.9
Model A2	−0.25	0.86
Model B1	−1.2	1.6
Model B2	−1.2	1.6

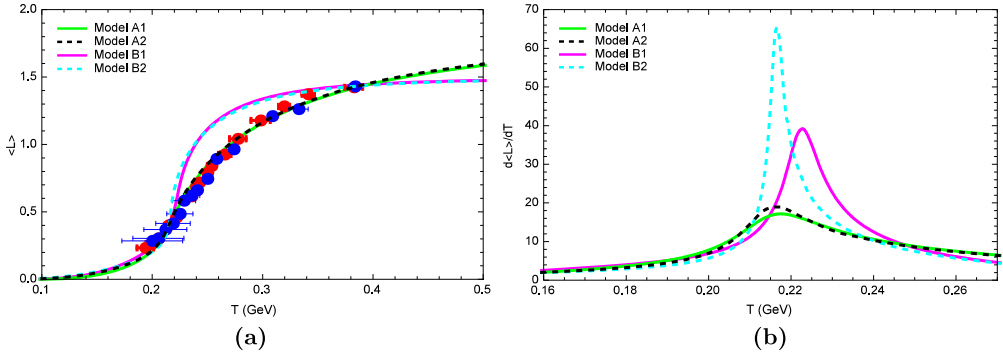


Fig. 7. The expectation value of the Polyakov loop $\langle L(T) \rangle$ (a) and its derivative $\frac{d\langle L \rangle}{dT}$ with respect to T (b) in model A1 (A2) and model B1 (B2). The red and blue points with error bars are two-flavor lattice data taken from [98]. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

in Table 2. One sees that the expectation value of Polyakov loop increases continuously with increasing temperature, which shows a crossover transition from confined phase to deconfined phase. In addition, we find again that the model A1 and A2 fit the lattice data much better than the model B1 and B2, which is consistent with the study of equation of state.

Since in the case of finite quark mass there is not a real phase transition but a crossover one, the transition temperature T_d is usually defined at the temperature where $\langle L \rangle$ changes fastest, i.e., the temperature with maximal $\frac{d\langle L \rangle}{dT}$. Therefore, we plot the derivative of $\langle L \rangle$ with respect to T in Fig. 7 (b), from which the pseudocritical temperature can be read from the location of the peak. We see that $T_d \simeq 217$ MeV in model A1, A2 and B2, while in Model B1 $T_d \simeq 223$ MeV. These results are consistent with the study of equation of state and comparable with the lattice results in [98].

4.2. Chiral phase transition from chiral condensate

In the previous sections, we have constrained our models by studying the equation of state and have tested these ones further in the Polyakov loop calculation. In this section, we will go further and try to investigate another important aspect of QCD phase transition, i.e., the chiral phase transition.

As mentioned above, chiral symmetry breaking and restoration are characterized by the order parameter, i.e., chiral condensate $\sigma \equiv \langle \bar{q}q \rangle$, which can be encoded in the scalar field X in the matter action (3) [79]. The matter action S_m has a $SU(2)_L \times SU(2)_R$ symmetry which is spontaneously broken when the scalar field X obtains a vacuum expectation value X_0 . Without loss of generality, we consider the case with degenerate quark mass, i.e., $m_u = m_d = m_q$, then we can

set $X_0(z) = \chi(z)I_2/2$ with I_2 the 2×2 identity matrix. As the chiral condensate σ only appears in $\chi(z)$, we derive the degenerate action of χ by inputting X_0 in the action (3) as follows

$$S_\chi = - \int d^5x \sqrt{-g} e^{-\phi} \left[\frac{1}{2} g^{zz} \chi'^2 + V(\chi) \right], \quad (38)$$

where $V(\chi) \equiv \text{Tr}(V_X(X))$. As noted in our previous work [83,84], the quartic term in the bulk scalar potential is necessary to realize the spontaneous chiral symmetry breaking. Hence, the potential $V(\chi)$ is given as

$$V(\chi) = -\frac{3}{2} \chi^2 + \lambda \chi^4. \quad (39)$$

Under the metric ansatz (7), the equation of motion for χ is easily derived from Eq. (38) as

$$\chi''(z) + \left(-\frac{3}{z} + 3A'_s(z) - \phi'(z) + \frac{f'(z)}{f(z)} \right) \chi'(z) - \frac{e^{2A_s(z)}}{z^2 f(z)} \partial_\chi V(\chi(z)) = 0, \quad (40)$$

from which the leading UV expansion of $\chi(z)$ can be obtained as follows

$$\chi(z) = m_q \zeta z + \frac{\sigma}{\zeta} z^3 + \dots, \quad (41)$$

where σ is the chiral condensate and the normalization constant $\zeta = \frac{\sqrt{3}}{2\pi}$ [23].

At first sight, m_q and σ are two independent integral constant of Eq. (40). Nevertheless, since $z = z_h$ is a singular point of Eq. (40), the regular condition of χ would require $\frac{1}{f(z)}(f' \chi' - e^{2A_s} \partial_\chi V(\chi)/z^2)$ to be finite at $z = z_h$, which means $f' \chi' - e^{2A_s} \partial_\chi V(\chi)/z^2 = 0$ at horizon. This would impose an IR boundary condition naturally, and cause the chiral condensate to depend on the quark mass and temperature (For details, see [83,84]). Using the above UV and IR boundary conditions, one can solve σ as a function of m_q and T .

4.2.1. Spontaneous chiral symmetry breaking and its restoration

In the chiral limit $m_q = 0$, there is always a trivial solution $\chi \equiv 0$ satisfying all the boundary conditions, which can be seen as the chiral symmetry restored phase. If there were chiral symmetry breaking in the system we studied, Eq. (40) would have non-trivial solutions with $\chi \neq 0$ at some temperature region. Using the background and the parameters fixed in the previous sections, we do find non-trivial solutions below certain temperature for all the four models, which shows the validity of the constraints we used. Here we only take model A1 as an example and show the results in Fig. 8 (a), in which we set $\lambda = 8$.

In order to study the thermodynamic stability of the solutions, we need to compare free energies of the different solutions and select the lowest free energy branch. According to the AdS/CFT dictionary, the free energy can be approximated by the on-shell action shown as follows [83,84]

$$\mathcal{F} \equiv \frac{F}{V_3} = - \left(\frac{1}{2z^3} \chi e^{3A_s - \phi} f \chi' \right) \Big|_\epsilon - \lambda \int_0^{z_h} dz \sqrt{-g} e^{-\phi} \chi^4. \quad (42)$$

In the chiral limit, the first term in the above expression vanishes. It is apparent that if $\lambda > 0$, the free energy of a non-trivial solution χ would be always smaller than the trivial solution $\chi = 0$. That means the physical vacuum has non-zero chiral condensate, and the spontaneous chiral symmetry breaking is realized. Furthermore, from Fig. 8 (a) one can see that the non-trivial

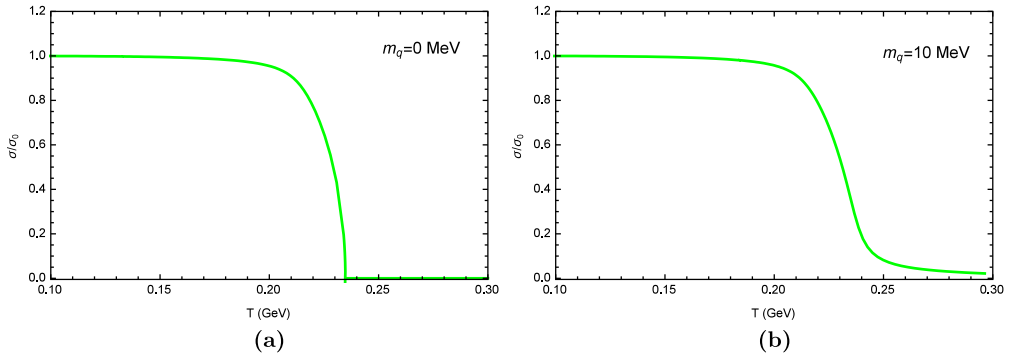


Fig. 8. The temperature dependence of chiral condensate σ in the chiral limit (a) and in finite quark mass case (b) in model A1.

Table 3

Chiral condensate related parameters in model A1 (A2) and model B1 (B2).

	m_q (MeV)	λ
Model A1	10	8
Model A2	20	8
Model B1	25	2
Model B2	25	2

solution disappears at around 235 MeV with only the trivial solution left, which means that at high temperature the chiral symmetry is restored. It is easy to read from the figure that the phase transition is of second order, which is consistent with the sketch plot shown in Fig. 1, since the derivative of σ with respect to T is discontinuous at the transition point.

Then we take finite m_q to calculate $\sigma(T)$ and find that the qualitative results are all similar to the one shown in Fig. 8 (b). However, when compared with Fig. 8 (a), we find that the trivial solution $\chi = 0$ has been disappeared as it does not satisfy the UV boundary condition (41). The chiral condensate decreases very slowly at low temperature and goes through a sudden drop at certain temperature, then goes to zero slowly, which shows obviously a crossover transition. Thus, under our background models, we obtain second order chiral phase transition in the chiral limit and crossover transition in the case of finite quark mass. This is qualitatively coincident with the sketch in Fig. 1.⁴

4.2.2. Confront chiral phase transition with lattice simulation

The lattice QCD studies of chiral phase transition in recent years have attracted much attention, especially in the $2 + 1$ flavor case with physical quark mass [100–103]. In [96,98], lattice simulation of thermal QCD transition with two light flavors was investigated for a set of pion masses ranging from 300 MeV to 600 MeV. Therefore, it is also interesting to compare our model calculation with those lattice results. We roughly fit the chiral condensate related parameters m_q , λ and present the results in Table 3.

⁴ Here it should be noted that this is an approximate result, since in the full analysis combining S_b and S_m , the quark mass would also affect the background.

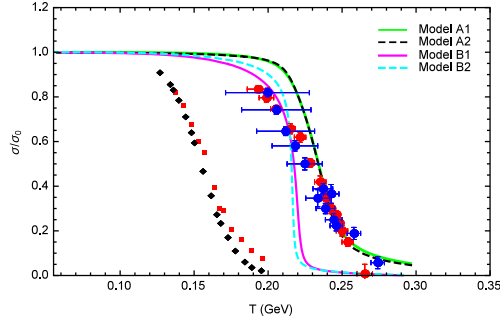


Fig. 9. Comparison of chiral phase transition in model A1 (A2) and model B1 (B2) with lattice results. The red and blue filled circles are lattice data with two light flavors in different lattice extent [98]. The red and black filled squares are lattice data with 2 + 1 flavors taken from [102]. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

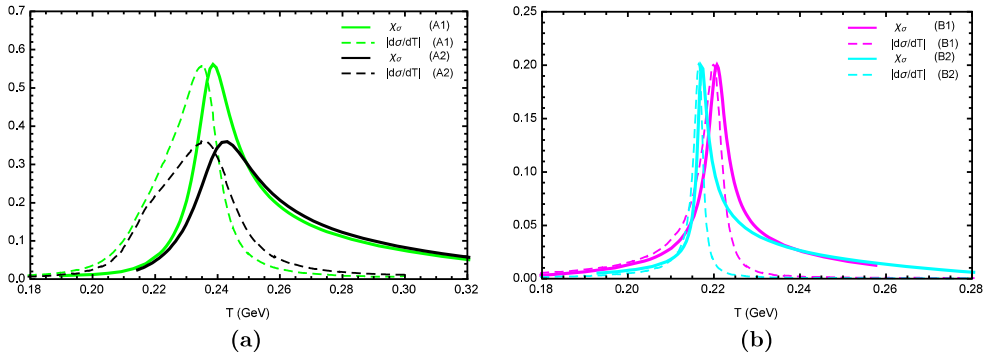


Fig. 10. Chiral susceptibility χ_σ and the derivative of condensate with respect to temperature $\frac{d\sigma}{dT}$ in model A1 (A2) (a) and model B1 (B2) (b).

The temperature dependence of the chiral condensate in our models are shown in Fig. 9. One can see that the behavior of chiral condensates with T in model A1 and A2 fit the lattice data better than that in model B1 and B2, which is consistent with the Polyakov loop and the equation of state calculations. The results of our models are much closer to the two-flavor lattice results than the ones of 2 + 1 flavor, which is also consistent with our previous calculations. From Fig. 9 one can also see that the chiral condensates decrease with temperature smoothly in the case of finite quark mass and go through a rapid drop when the temperature reaches to a value which characterizes the crossover transition of thermal QCD. To fix the chiral transition temperature T_c , we calculate the chiral susceptibility which is defined as

$$\chi_\sigma = \frac{\partial \langle \bar{\psi} \psi \rangle}{\partial m_q} = \frac{\partial \sigma}{\partial m_q}. \quad (43)$$

As a crosscheck, we also give another criterion to determine chiral transition temperature by the quantity $\frac{d\sigma}{dT}$. The pseudocritical temperature is identified as the extremal point of χ_σ or $\frac{d\sigma}{dT}$. The numerical results are presented in Fig. 10. One can see that although the curves of χ_σ and $\frac{d\sigma}{dT}$ are different overall, their extremal points lie in a narrow temperature region which indicates the possible transition temperature. The pseudocritical temperatures in model A1 and

A2 are around $235 \sim 242$ MeV, while the ones in model B1 and B2 are about $217 \sim 221$ MeV. Compared with the previous results, one can see that the chiral transition temperature T_c and the deconfining transition temperature T_d in model B1 and B2 are much close with each other, while in model A1 and A2 T_c seems a little higher than T_d . Fixing the two transition temperatures both from experimental side and from theoretical side is an interesting issue of many years and there are still many debates and discrepancies. We just show our results for a possible case.

5. Discussion and conclusion

Thermal QCD studies of chiral and deconfining phase transitions have attracted much attention both from Lattice QCD [96–98,102–105] and from effective field theories such as NJL models [106–108]. In the holographic framework, the crossover behavior of the two phase transitions at physical quark mass has not yet been realized simultaneously. In this paper, we make the first attempt to accommodate these two phase transitions in a single holographic QCD model.

We consider the Einstein-dilaton system coupled with the soft-wall model action. As a preliminary try, we ignore the back-reaction from the flavor part to the background geometry. The flavor effects are taken into account by comparing the equation of state generating from the background part with those of two-flavor lattice simulation. By analyzing the Einstein-dilaton system, together with the constraints from the qualitative behavior of the two phase transitions, we propose four types of bottom-up holographic models (A1, A2, B1, B2). Then we use potential reconstruction approach to solve the whole background system. We find that within the physically concerned region, the temperature dependence of the dilaton potential is very weak, which imply the validity of the method we used. Under these backgrounds, we calculate the equation of state and compare with the two-flavor lattice QCD results. We find that our results of pressure, energy density and trace anomaly are in good agreement with the two-flavor lattice results in [96]. We also see the sudden release of degrees of freedom from the entropy density, energy density and pressure, which indicate the crossover behavior of thermal transition.

Then we try to study temperature dependent behavior of Polyakov loop and chiral condensate, which are the corresponding order parameters of deconfining and chiral phase transition respectively. We find that all the four models give the correct qualitative behavior of Polyakov loop. All of them show a crossover transition from confined phase to deconfined phase with increasing temperature. Furthermore, the calculations in models A1 and A2 fit the two-flavor lattice results from [98] in very good accuracy. After that, we investigate chiral condensate in the matter part. We find that in the chiral limit, the spontaneous chiral symmetry breaking and its restoration are correctly realized. And then we take finite quark mass and compare the results with the two-flavor lattice ones from [98]. We find crossover behavior from a chiral symmetry breaking phase to a restored phase in the chiral transition in all the four models.

The transition temperature extracted from the location of the valley of sound speed square is around 217 MeV for models A1, A2 and B2, while for model B1 is around 223 MeV. We note that this is very close to the one extracted from the peak of the derivative of Polyakov loop with respect to temperature, which is reasonable since the valley of sound speed square contains the information of degree of freedom in both phases and the deconfining phase transition is related to the release of color degrees of freedom. This can be seen as one of the consistency checks in our models. We also extract the transition temperature from the chiral susceptibility χ_σ and the derivative of chiral condensate with respect to temperature. We note that the deviation of the transition temperatures extracted from the two ways is small, so a reasonable chiral transition temperature can be obtained. For the models A1 and A2 the transition temperatures are both

around $235 \sim 242$ MeV, which is slightly larger than the deconfining temperature, while for the models B1 and B2 they are around $217 \sim 221$ MeV, which is very close to the corresponding deconfining temperature.

Finally, we emphasize that the studies here are very preliminary, and the main motivation here is to grasp the necessary ingredients in characterizing the two phase transitions simultaneously. We have not considered the back-reaction of the flavor part to the background geometry. We also note that the models A1 and A2 give better fittings to the lattice results in all the quantities, including the equation of state, the Polyakov loop and the chiral condensate. In this sense, the dominant effects in the intermediate region might be the z^2 term rather than z^4 . These studies could provide clues for the more careful research in the future. It is also quite interesting to study the chemical potential effects in these models, which would be finished in the near future.

Acknowledgements

We are grateful to Ronggen Cai, Mei Huang, Shigeki Sugimoto, Tadashi Takayanagi, Yue-Liang Wu and Yi Yang for useful conversations and correspondence. Z.F. thanks Yue-Liang Wu for his warm guidance and patient instructions. S.H. thanks Tadashi Takayanagi for their encouragement and support. S.H. is supported by JSPS postdoctoral fellowship for foreign researchers and by the National Natural Science Foundation of China (No. 11305235). This work is funded in part by China Postdoctoral Science Foundation (No. 2015M580136).

References

- [1] Y. Aoki, G. Endrodi, Z. Fodor, S.D. Katz, K.K. Szabo, *Nature* 443 (2006) 675, arXiv:hep-lat/0611014.
- [2] E. Laermann, O. Philipsen, *Annu. Rev. Nucl. Part. Sci.* 53 (2003) 163, arXiv:hep-ph/0303042.
- [3] P. de Forcrand, O. Philipsen, *J. High Energy Phys.* 0701 (2007) 077, arXiv:hep-lat/0607017.
- [4] K. Kanaya, *AIP Conf. Proc.* 1343 (2011) 57, arXiv:1012.4235 [hep-ph].
- [5] R.D. Pisarski, F. Wilczek, *Phys. Rev. D* 29 (1984) 338.
- [6] C.S. Fischer, J.A. Mueller, *Phys. Rev. D* 80 (2009) 074029, arXiv:0908.0007 [hep-ph].
- [7] J.O. Andersen, W.R. Naylor, A. Tranberg, *J. High Energy Phys.* 1404 (2014) 187, arXiv:1311.2093 [hep-ph].
- [8] C.S. Fischer, J. Luecker, J.A. Mueller, *Phys. Lett. B* 702 (2011) 438, arXiv:1104.1564 [hep-ph].
- [9] F. Xu, H. Mao, T.K. Mukherjee, M. Huang, *Phys. Rev. D* 84 (2011) 074009, arXiv:1104.0873 [hep-ph].
- [10] J.M. Maldacena, *Adv. Theor. Math. Phys.* 2 (1998) 231, *Int. J. Theor. Phys.* 38 (1999) 1113, arXiv:hep-th/9711200.
- [11] S.S. Gubser, I.R. Klebanov, A.M. Polyakov, *Phys. Lett. B* 428 (1998) 105, arXiv:hep-th/9802109.
- [12] E. Witten, *Adv. Theor. Math. Phys.* 2 (1998) 253, arXiv:hep-th/9802150.
- [13] J. Erlich, E. Katz, D.T. Son, M.A. Stephanov, *Phys. Rev. Lett.* 95 (2005) 261602, arXiv:hep-ph/0501128.
- [14] G.F. de Teramond, S.J. Brodsky, *Phys. Rev. Lett.* 94 (2005) 201601, arXiv:hep-th/0501022.
- [15] L. Da Rold, A. Pomarol, *Nucl. Phys. B* 721 (2005) 79, arXiv:hep-ph/0501218.
- [16] J. Babington, J. Erdmenger, N.J. Evans, Z. Guralnik, I. Kirsch, *Phys. Rev. D* 69 (2004) 066007, arXiv:hep-th/0306018.
- [17] M. Kruczenski, D. Mateos, R.C. Myers, D.J. Winters, *J. High Energy Phys.* 0405 (2004) 041, arXiv:hep-th/0311270.
- [18] T. Sakai, S. Sugimoto, *Prog. Theor. Phys.* 113 (2005) 843, arXiv:hep-th/0412141.
- [19] T. Sakai, S. Sugimoto, *Prog. Theor. Phys.* 114 (2006) 1083, arXiv:hep-th/0507073.
- [20] C. Csaki, M. Reece, J. High Energy Phys. 0705 (2007) 062, arXiv:hep-ph/0608266.
- [21] S. He, M. Huang, Q.S. Yan, Y. Yang, *Eur. Phys. J. C* 66 (2010) 187, arXiv:0710.0988 [hep-ph].
- [22] T. Gherghetta, J.I. Kapusta, T.M. Kelley, *Phys. Rev. D* 79 (2009) 076003, arXiv:0902.1998 [hep-ph].
- [23] A. Cherman, T.D. Cohen, E.S. Werbos, *Phys. Rev. C* 79 (2009) 045203, arXiv:0804.1096 [hep-ph].
- [24] T.M. Kelley, S.P. Bartz, J.I. Kapusta, *Phys. Rev. D* 83 (2011) 016002, arXiv:1009.3009 [hep-ph].
- [25] Y.-Q. Sui, Y.-L. Wu, Z.-F. Xie, Y.-B. Yang, *Phys. Rev. D* 81 (2010) 014024, arXiv:0909.3887 [hep-ph].
- [26] Y.-Q. Sui, Y.-L. Wu, Y.-B. Yang, *Phys. Rev. D* 83 (2011) 065030, arXiv:1012.3518 [hep-ph].
- [27] L.X. Cui, Z. Fang, Y.L. Wu, *Eur. Phys. J. C* 76 (1) (2016) 22.

- [28] D. Li, M. Huang, Q.S. Yan, Eur. Phys. J. C 73 (2013) 2615, arXiv:1206.2824 [hep-th].
- [29] D. Li, M. Huang, J. High Energy Phys. 1311 (2013) 088, arXiv:1303.6929 [hep-ph].
- [30] E.V. Shuryak, Nucl. Phys. A 750 (2005) 64, arXiv:hep-ph/0405066.
- [31] M.J. Tannenbaum, Rep. Prog. Phys. 69 (2006) 2005, arXiv:nucl-ex/0603003.
- [32] G. Policastro, D.T. Son, A.O. Starinets, Phys. Rev. Lett. 87 (2001) 081601, arXiv:hep-th/0104066.
- [33] R.-G. Cai, Z.-Y. Nie, N. Ohta, Y.-W. Sun, Phys. Rev. D 79 (2009) 066004, arXiv:0901.1421 [hep-th].
- [34] R.-G. Cai, Z.-Y. Nie, Y.-W. Sun, Phys. Rev. D 78 (2008) 126007, arXiv:0811.1665 [hep-th].
- [35] S.J. Sin, I. Zahed, Phys. Lett. B 608 (2005) 265, arXiv:hep-th/0407215.
- [36] E. Shuryak, S.J. Sin, I. Zahed, J. Korean Phys. Soc. 50 (2007) 384, arXiv:hep-th/0511199.
- [37] H. Nastase, arXiv:hep-th/0501068.
- [38] R.A. Janik, R.B. Peschanski, Phys. Rev. D 73 (2006) 045013, arXiv:hep-th/0512162.
- [39] S. Nakamura, S.J. Sin, J. High Energy Phys. 0609 (2006) 020, arXiv:hep-th/0607123.
- [40] S.J. Sin, S. Nakamura, S.P. Kim, J. High Energy Phys. 0612 (2006) 075, arXiv:hep-th/0610113.
- [41] C.P. Herzog, A. Karch, P. Kovtun, C. Kozcaz, L.G. Yaffe, J. High Energy Phys. 0607 (2006) 013, arXiv:hep-th/0605158.
- [42] S.S. Gubser, Phys. Rev. D 74 (2006) 126005, arXiv:hep-th/0605182.
- [43] Z.q. Zhang, D.f. Hou, H.c. Ren, J. High Energy Phys. 1301 (2013) 032, arXiv:1210.5187 [hep-th].
- [44] Y. Wu, D. Hou, H.c. Ren, arXiv:1401.3635 [hep-ph].
- [45] D. Li, J. Liao, M. Huang, Phys. Rev. D 89 (12) (2014) 126006, arXiv:1401.2035 [hep-ph].
- [46] D. Li, S. He, M. Huang, J. High Energy Phys. 1506 (2015) 046, arXiv:1411.5332 [hep-ph].
- [47] O. Aharony, S.S. Gubser, J.M. Maldacena, H. Ooguri, Y. Oz, Phys. Rep. 323 (2000) 183, arXiv:hep-th/9905111.
- [48] J. Erdmenger, N. Evans, I. Kirsch, E. Threlfall, Eur. Phys. J. A 35 (2008) 81, arXiv:0711.4467 [hep-th].
- [49] S.J. Brodsky, G.F. de Teramond, H.G. Dosch, J. Erlich, Phys. Rep. 584 (2015) 1, arXiv:1407.8131 [hep-ph].
- [50] Y. Kim, I.J. Shin, T. Tsukioka, Prog. Part. Nucl. Phys. 68 (2013) 55, arXiv:1205.4852 [hep-ph].
- [51] A. Adams, L.D. Carr, T. Schäfer, P. Steinberg, J.E. Thomas, New J. Phys. 14 (2012) 115009, arXiv:1205.5180 [hep-th].
- [52] C.P. Herzog, Phys. Rev. Lett. 98 (2007) 091601, arXiv:hep-th/0608151.
- [53] C.A. Ballon Bayona, H. Boschi-Filho, N.R.F. Braga, L.A. Pando Zayas, Phys. Rev. D 77 (2008) 046002, arXiv:0705.1529 [hep-th].
- [54] R.G. Cai, J.P. Shock, J. High Energy Phys. 0708 (2007) 095, arXiv:0705.3388 [hep-th].
- [55] R.G. Cai, S. Chakraborty, S. He, L. Li, J. High Energy Phys. 1302 (2013) 068, arXiv:1209.4512 [hep-th].
- [56] Y. Kim, B.H. Lee, S. Nam, C. Park, S.J. Sin, Phys. Rev. D 76 (2007) 086003, arXiv:0706.2525 [hep-ph].
- [57] Y. Kim, T. Misumi, I.J. Shin, arXiv:0911.3205 [hep-ph].
- [58] O. Andreev, Phys. Rev. Lett. 102 (2009) 212001, arXiv:0903.4375 [hep-ph].
- [59] P. Colangelo, F. Giannuzzi, S. Nicotri, Phys. Rev. D 83 (2011) 035015, arXiv:1008.3116 [hep-ph].
- [60] S.S. Gubser, A. Nellore, S.S. Pufu, F.D. Rocha, Phys. Rev. Lett. 101 (2008) 131601, arXiv:0804.1950 [hep-th].
- [61] S.S. Gubser, A. Nellore, Phys. Rev. D 78 (2008) 086007, arXiv:0804.0434 [hep-th].
- [62] S.S. Gubser, S.S. Pufu, F.D. Rocha, J. High Energy Phys. 0808 (2008) 085, arXiv:0806.0407 [hep-th].
- [63] U. Gursoy, E. Kiritsis, L. Mazzanti, F. Nitti, Phys. Rev. Lett. 101 (2008) 181601, arXiv:0804.0899 [hep-th].
- [64] U. Gursoy, E. Kiritsis, J. High Energy Phys. 0802 (2008) 032, arXiv:0707.1324 [hep-th].
- [65] U. Gursoy, E. Kiritsis, F. Nitti, J. High Energy Phys. 0802 (2008) 019, arXiv:0707.1349 [hep-th].
- [66] U. Gursoy, E. Kiritsis, L. Mazzanti, F. Nitti, J. High Energy Phys. 0905 (2009) 033, arXiv:0812.0792 [hep-th].
- [67] S.I. Finazzo, J. Noronha, Phys. Rev. D 90 (11) (2014) 115028, arXiv:1411.4330 [hep-th].
- [68] R. Yaresko, B. Kampfer, Phys. Lett. B 747 (2015) 36, arXiv:1306.0214 [hep-ph].
- [69] R. Yaresko, J. Knaute, B. Kampfer, Eur. Phys. J. C 75 (6) (2015) 295, arXiv:1503.09065 [hep-ph].
- [70] D. Li, S. He, M. Huang, Q.S. Yan, J. High Energy Phys. 1109 (2011) 041, arXiv:1103.5389 [hep-th].
- [71] R.G. Cai, S. He, D. Li, J. High Energy Phys. 1203 (2012) 033, arXiv:1201.0820 [hep-th].
- [72] S. He, S.Y. Wu, Y. Yang, P.H. Yuan, J. High Energy Phys. 1304 (2013) 093, arXiv:1301.0385 [hep-th].
- [73] Y. Yang, P.H. Yuan, J. High Energy Phys. 1411 (2014) 149, arXiv:1406.1865 [hep-th].
- [74] Y. Yang, P.H. Yuan, arXiv:1506.05930 [hep-th].
- [75] F. Zuo, J. High Energy Phys. 1406 (2014) 143, arXiv:1404.4512 [hep-ph].
- [76] F. Zuo, Y.H. Gao, J. High Energy Phys. 1407 (2014) 147, arXiv:1403.2241 [hep-ph].
- [77] L.X. Cui, Z. Fang, Y.L. Wu, arXiv:1404.0761 [hep-ph].
- [78] S. Afonin, A.D. Katanaeva, Eur. Phys. J. C 74 (10) (2014) 3124, arXiv:1408.6935 [hep-ph].
- [79] A. Karch, E. Katz, D.T. Son, M.A. Stephanov, Phys. Rev. D 74 (2006) 015005, arXiv:hep-ph/0602229.
- [80] Y. Nambu, G. Jona-Lasinio, Phys. Rev. 122 (1961) 345.

- [81] Y. Nambu, G. Jona-Lasinio, Phys. Rev. 124 (1961) 246.
- [82] P. Colangelo, F. Giannuzzi, S. Nicotri, V. Tangorra, Eur. Phys. J. C 72 (2012) 2096, arXiv:1112.4402 [hep-ph].
- [83] K. Chelabi, Z. Fang, M. Huang, D. Li, Y.L. Wu, arXiv:1511.02721 [hep-ph].
- [84] K. Chelabi, Z. Fang, M. Huang, D. Li, Y.L. Wu, arXiv:1512.06493 [hep-ph].
- [85] C. Nunez, A. Paredes, A.V. Ramallo, Adv. High Energy Phys. 2010 (2010) 196714, arXiv:1002.1088 [hep-th].
- [86] M. Jarvinen, E. Kiritsis, J. High Energy Phys. 1203 (2012) 002, arXiv:1112.1261 [hep-ph].
- [87] S. He, Y.-P. Hu, J.-H. Zhang, J. High Energy Phys. 1112 (2011) 078, arXiv:1111.1374 [hep-th].
- [88] S. He, M. Huang, Q.-S. Yan, Phys. Rev. D 83 (2011) 045034, arXiv:1004.1880 [hep-ph].
- [89] K. Kajantie, M. Krssak, M. Vepsalainen, A. Vuorinen, Phys. Rev. D 84 (2011) 086004, arXiv:1104.5352 [hep-ph].
- [90] B. Batell, T. Gherghetta, Phys. Rev. D 78 (2008) 026002, arXiv:0801.4383 [hep-ph].
- [91] F.V. Gubarev, L. Stodolsky, V.I. Zakharov, Phys. Rev. Lett. 86 (2001) 2220–2222, arXiv:hep-ph/0010057.
- [92] F.V. Gubarev, V.I. Zakharov, Phys. Lett. B 501 (2001) 28–36, arXiv:hep-ph/0010096.
- [93] K.I. Kondo, Phys. Lett. B 514 (2001) 335, arXiv:hep-th/0105299.
- [94] D. Zwanziger, Nucl. Phys. B 345 (1990) 461.
- [95] E. Megias, H.J. Pirner, K. Veschgini, Phys. Rev. D 83 (2011) 056003, arXiv:1009.2953 [hep-ph].
- [96] F. Burger, et al., tmfT Collaboration, Phys. Rev. D 91 (7) (2015) 074504, arXiv:1412.6748 [hep-lat].
- [97] A. Bazavov, et al., HotQCD Collaboration, Phys. Rev. D 90 (9) (2014) 094503, arXiv:1407.6387 [hep-lat].
- [98] F. Burger, et al., tmfT Collaboration, Phys. Rev. D 87 (7) (2013) 074508, arXiv:1102.4530 [hep-lat].
- [99] J. Noronha, Phys. Rev. D 81 (2010) 045011, arXiv:0910.1261 [hep-th].
- [100] S. Borsanyi, et al., Phys. Rev. D 92 (1) (2015) 014505, arXiv:1504.03676 [hep-lat].
- [101] Y. Aoki, S. Borsanyi, S. Durr, Z. Fodor, S.D. Katz, S. Krieg, K.K. Szabo, J. High Energy Phys. 0906 (2009) 088, arXiv:0903.4155 [hep-lat].
- [102] A. Bazavov, et al., Phys. Rev. D 85 (2012) 054503, arXiv:1111.1710 [hep-lat].
- [103] T. Bhattacharya, et al., Phys. Rev. Lett. 113 (8) (2014) 082001, arXiv:1402.5175 [hep-lat].
- [104] S. Borsanyi, Z. Fodor, C. Hoelbling, S.D. Katz, S. Krieg, K.K. Szabo, Phys. Lett. B 730 (2014) 99, arXiv:1309.5258 [hep-lat].
- [105] V.G. Bornyakov, R. Horsley, S.M. Morozov, Y. Nakamura, M.I. Polikarpov, P.E.L. Rakow, G. Schierholz, T. Suzuki, Phys. Rev. D 82 (2010) 014504, arXiv:0910.2392 [hep-lat].
- [106] C. Ratti, M.A. Thaler, W. Weise, Phys. Rev. D 73 (2006) 014019, arXiv:hep-ph/0506234.
- [107] S. Roessner, C. Ratti, W. Weise, Phys. Rev. D 75 (2007) 034007, arXiv:hep-ph/0609281.
- [108] S. Roessner, T. Hell, C. Ratti, W. Weise, Nucl. Phys. A 814 (2008) 118, arXiv:0712.3152 [hep-ph].