## **SIMPLE limits on spin-dependent WIMP interactions**

TA Girard<sup>a</sup>, F Giuliani<sup>a</sup>, T Morlat<sup>a</sup>, JI Collar<sup>b</sup>, D Limagne<sup>c</sup>, G Waysand<sup>c,d</sup>, J Puibasset<sup>c</sup>, HS Miley<sup>c</sup>, M Auguste<sup>d</sup>, D Boyer<sup>d</sup>, A Cavaillou<sup>d</sup>, JG Marques<sup>f,a</sup>, C Oliveira<sup>f</sup>, AC Fernandes<sup>f,a</sup>, AR Ramos<sup>f,a</sup>, M Felizardo<sup>f,g</sup> and RC Martins<sup>g</sup>

<sup>a</sup> Centro de Física Nuclear, Universidade de Lisboa, 1649-003 Lisbon, Portugal

<sup>b</sup> Department of Physics, University of Chicago, Chicago IL, 60637 USA

<sup>e</sup> Pacific Northwest National Laboratory, Richland, WA 99352 USA

<sup>d</sup> Laboratoire Souterrain à Bas Bruit, 84400 Rustrel-Pays d'Apt, France

<sup>f</sup> Instituto Tecnológico e Nuclear, 2686-953 Sacavém, Portugal

<sup>g</sup> Department of Electronics, Instituto Superior Técnico, 1049-001 Lisbon, Portugal

E-mail: criodets@cii.fc.ul.pt

**Abstract**. An improved SIMPLE experiment comprising four superheated droplet detectors with a total exposure of 0.42 kg.d yields ~factor 10 improvement in the previously-reported results. Despite the low exposure, the result provides restrictions on the allowed phase space of spin-dependent coupling strengths almost equivalent to those from the significantly larger exposure NAIAD-CDMS/ZEPLIN searches.

SIMPLE is a project to detect spin-sensitive WIMP dark matter interactions, and one of only two [1] such projects based on superheated droplet detectors (SDDs). The basics of these devices have been described previously [2]. We report the results of our most recent measurement [3].

The detectors were fabricated in-house from  $C_2ClF_5$  (R-115) according to previously-described procedures [2]. These were installed in the 1500 mwe LSBB laboratory [4], inside a thermally-regulated 700 liter water bath, surrounded by three layers of sound and thermal insulation, resting on a dual vibration absorber. To reduce ambient noise, a hydrophone was placed within the detector water bath, and a second acoustic monitor positioned outside the shielding. Apart from temperature control, the surrounding water bath acts as a ~ 30 cm thick neutron moderator, further reducing any ambient neutron flux by at least two orders of magnitude.

A bubble nucleation is detected by a piezoelectric transducer immersed in a glycerine layer at the top of the detector. Whenever an event in any one of the detectors occurs, the temperature, pressure, and threshold voltage level for each device, plus its waveform trace and fast Fourier transform, are recorded in a Labview platform.

The detectors were operated for 10.2 days at 8.9°C (2.0 atm), and 14.3 days at 3.3°C (1.9 atm). Additional measurements were performed at 14°C in order to insure that all low rate devices were operating properly.

In a first analysis stage, the individual detectors were physically inspected for fractures (which lead to spontaneous nucleations at the fracture sites), and their responses monitored over the measurement period. Three of the devices were rejected as a result of fractures and/or performances outside specification tolerances.

<sup>&</sup>lt;sup>c</sup> INSP (UMR 7588 CNRS), Universités Paris 7 & 6, 75251 Paris, France

The data record was anti-coincidence filtered on an event-by-event basis, with the criteria that (i) one and only one of the in-bath detectors had a signal, and (ii) no monitoring detector had a simultaneous signal.

			Filter N°1		Filter N°2	
Detector	Active	Acoustic	8.9°C	3.3°C	8.9°C	3.3°C
	Mass (g)	Efficiency	(evts/kg.d)	(evts/kg.d)	(evts/kg.d)	(evts/kg.d)
2	9.9	$0.52\pm0.10$	$278\pm52.6$	$21.2\pm12.2$	$179 \pm 42.1$	$7.1 \pm 7.1$
4	10.8	$1.11\pm0.19$	$54.6 \pm 22.3$	$25.9 \pm 13.0$	$36.4 \pm 18.2$	$25.9 \pm 13.0$
5	10.4	$0.83\pm0.17$	$66.2\pm25.0$	$13.4 \pm 9.5$	$28.4 \pm 16.4$	0
7	11.1	$1.21\pm0.21$	$558\pm70.3$	$271 \pm 41.3$	$407\pm60.1$	$221\pm37.3$

Table 1. Data results, without acoustic detection efficiency or background correction.

A second filtering was imposed in which only the Filter N°1 events with a primary harmonic between 5.5-6.5 kHz were accepted, since the frequency spectrum of the transducer signal comprises a well-defined response with a primary harmonic at ~ 6 kHz. The rate reductions are shown in Table 1.

Backgrounds and their treatment have been discussed elsewhere [3]. The difference (*n*) between the fully filtered measurements at 9°C and 3°C is most likely a sum of background and microleak events: a 90% C.L. upper limit to the unobserved WIMP rate was set from the expectation value of the total number of events,  $\mu_n$ , such that the probability of observing at least *n* events is 90%. Subtracting the expected number of background events (computed by maximum likelihood) from  $\mu_n$  yields an estimate of 24.0 events/(kg.d) (corrected for acoustic detection efficiency) for the expected, unobserved WIMP events.



**Figure 1.**  $a_p - a_n$  for SIMPLE (thick dashed), PICASSO (thick solid), NAIAD (dotted), CRESST-I and Tokyo/NaF for WIMP mass of 50 GeV/c<sup>2</sup>. Also shown are the single nuclei DAMA/Xe2, and spin-INdependent EDELWEISS, CDMS and ZEPLIN-I. The region permitted by each experiment is the area inside the respective contour, with the shaded central region the allowed intersection of the NAIAD-CDMS/ZEPLIN measurements [3].

The results were analyzed in a model-independent formalism following reference [5] using the cosmological parameters and method described in reference [6], in which the spin-dependent

interaction cross section for a WIMP interaction with a nucleon is  $\sigma_{sD} \sim [a_p < S_p > + a_n < S_n >]^2$ , where  $a_{p,n}$  ( $S_{p,n}$ ) are the proton and neutron coupling strengths (proton and neutron group spins) respectively. These are shown in figure 1 at 90% C.L. for  $M_w = 50 \text{ GeV/c}^2$  (masses above or below this choice yield slightly increased limits), using the spin values of reference [7]; use of the reference [8] values would rotate the SIMPLE and PICASSO curves about the origin to a more horizontal orientation.

The region excluded by an experiment lies outside the indicated band, and the allowed region is defined by the intersection of the various bands. The new SIMPLE result eliminates a large part of the parameter space allowed by the significantly larger exposure Tokyo/NaF [9], NAIAD/NaI [10] and CRESST-I/Al<sub>2</sub>O<sub>3</sub> [11] measurements at this mass cut, as well as the neutron-sensitive DAMA/Xe2 experiment [12]. The SIMPLE result provides slightly more restrictive bounds than the most recent 2 kg.d PICASSO report [1], but with a factor ~5 less exposure; the difference most likely results from the higher intrinsic backgrounds owing to the CsCl salts used in density-matching their gel and refrigerant. The allowed area of the SIMPLE-CDMS/ZEPLIN intersection ( $|a_p| \le 2.4$ ,  $|a_n| \le 0.8$ ) at 50 GeV/c<sup>2</sup> is only slightly larger than that of NAIAD-CDMS/ZEPLIN ( $|a_p| \le 1.4$ ,  $|a_n| \le 0.7$ ), and corresponds to [5]  $\sigma_a \le 0.7$  pb,  $\sigma_a \le 0.2$  pb.

Penetrating the frontier of the current allowed region of the  $a_p - a_n$  phase space shown in figure 1 would require only a modest 3 kg.d exposure at the current SIMPLE performance, which is relatively easily achievable. The recent application of pulse shape analysis techniques to the data records has moreover identified the possibility of discriminating between signal and microleak events, suggesting the ability to reach overall measurement rates of ~ 1 event/(kg.d) for the same 3 kg.d exposure, corresponding to a factor of 10 further reduction in the exclusion. The SIMPLE project has recently received funding for conduct of the 3 kg.d measurement.

## Acknowledgments

This work was supported in part by grants POCTI FP/57834/2004 and FNU/43683/2002 of the Foundation for Science and Technology of Portugal, co-financed by FEDER.

## References

- [1] Barnabé-Heider M et al 2005 Phys. Lett. B 624 186
- [2] Collar JI, Puibasset J, Girard TA et al 2000 Phys. Rev. Lett. 85 3083
- [3] Girard TA, Giuliani F, Collar JI et al 2005 Phys. Lett. B 621 233
- [4] Laboratoire Souterrain à Bas Bruit de Rustrel-Pays d'Apt: http://lsbb.unice.fr.
- [5] Giuliani F 2004 *Phys. Rev. Lett.* 93 161301
  Giuliani F and Girard TA 2005 *Phys. Rev.* D 71 123503
- [6] Lewin JD and Smith PF 1996 Astrop. Phys. 6 87
- [7] Pacheco F and Strottman D 1989 Phys. Rev. D 40 2131
- [8] Divari MT, Kosmas TS, Vergados JD et al 2000 Phys. Rev. C 61 054612
- [9] Takeda A, Minowa M, Miuchi K et al 2003 Phys. Lett. B 572 145
- [10] Ahmed B, Alner GJ, Araujo H et al 2003 Astrop. Phys. 19 691
- [11] Seidel W et al 2002 Dark Matter in Astro- and Particle Physics (Berlin: Springer)
- [12] Bernabei R, Belli P, Montecchia F et al 1998 Phys. Lett. B 436 379